

# Transient phenomena in foreshock, shock, and magnetosheath – Expectations from large separation campaign

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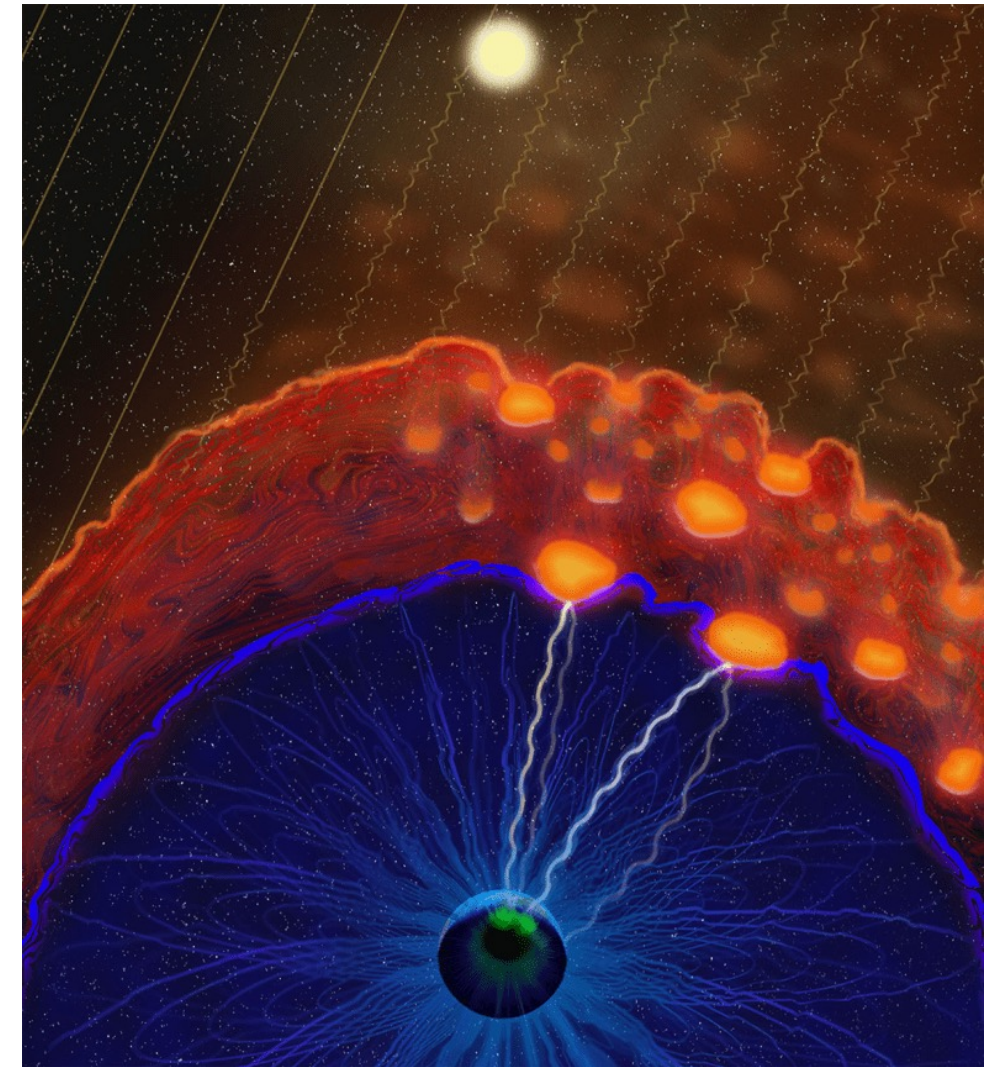
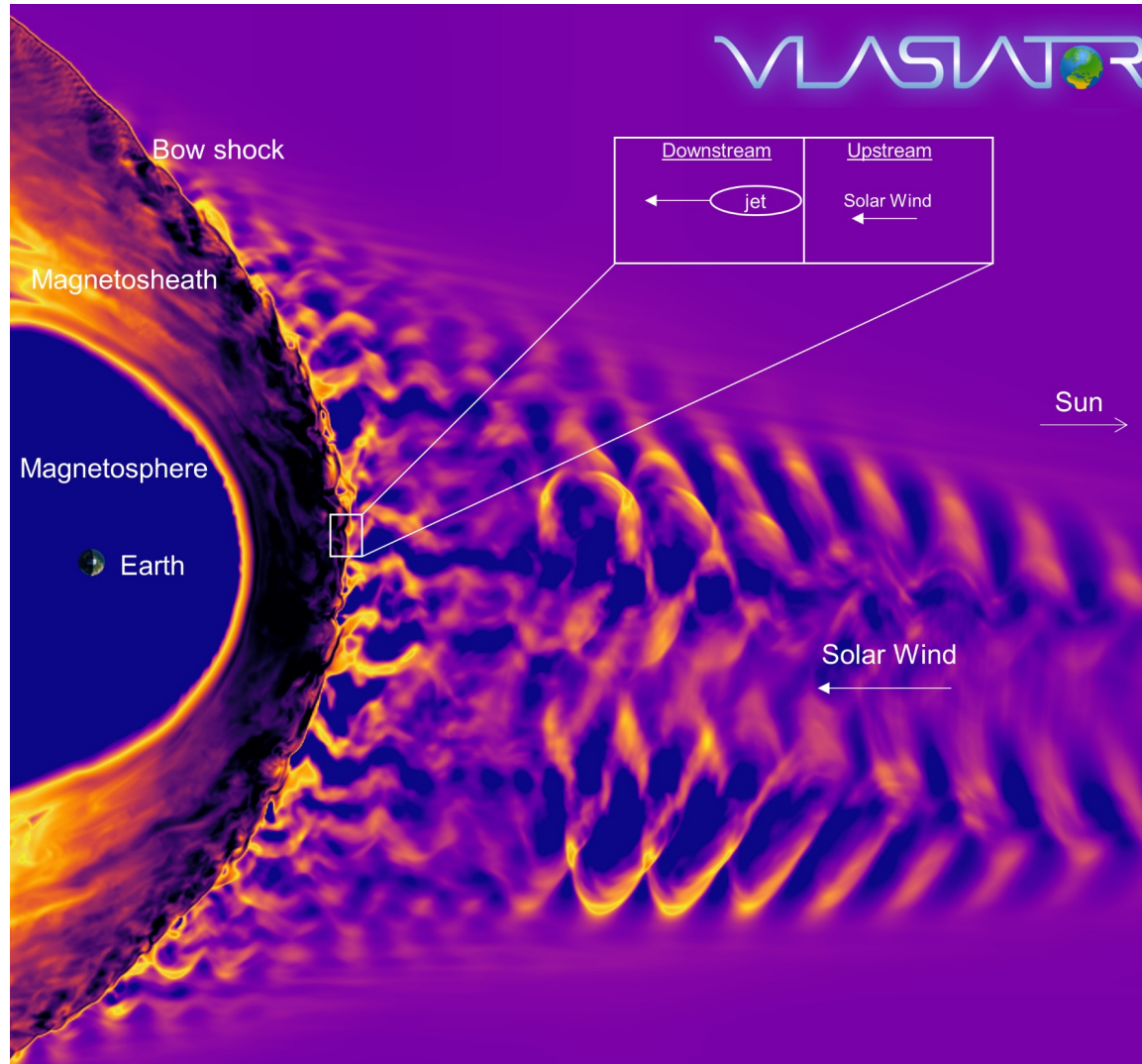
SR acknowledges the support by John Hopkins University Applied Physics Laboratory  
independent R&D fund

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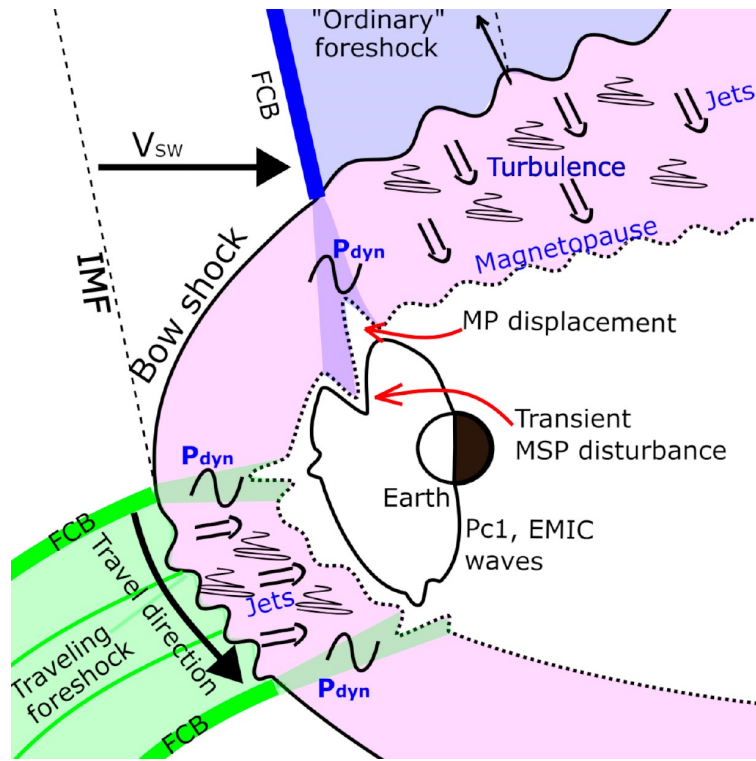


CENTER FOR  
GEOSPACE STORMS

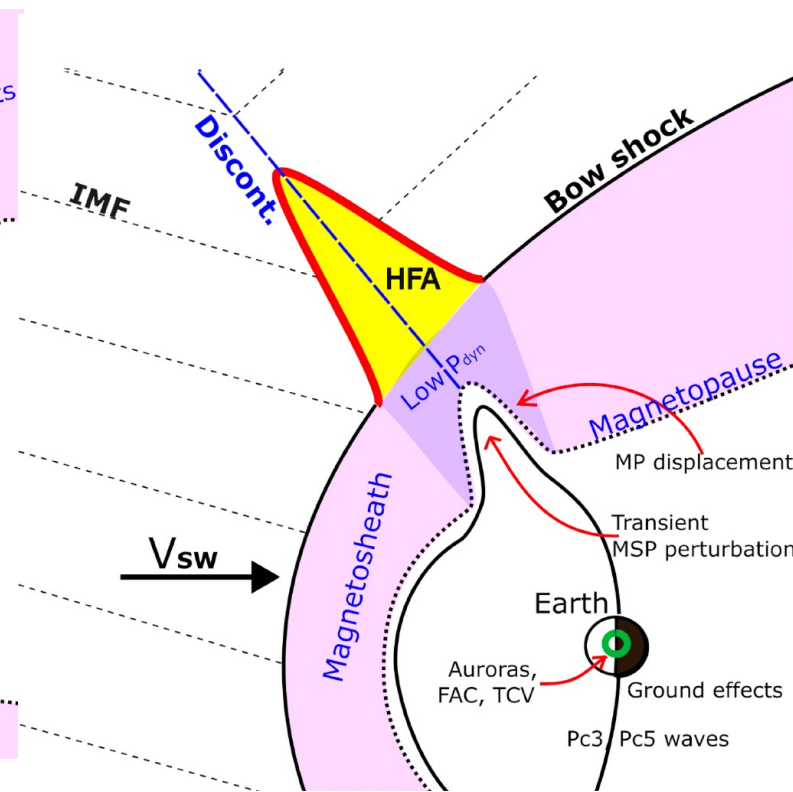
# Transient Phenomena in Foreshock & Magnetosheath



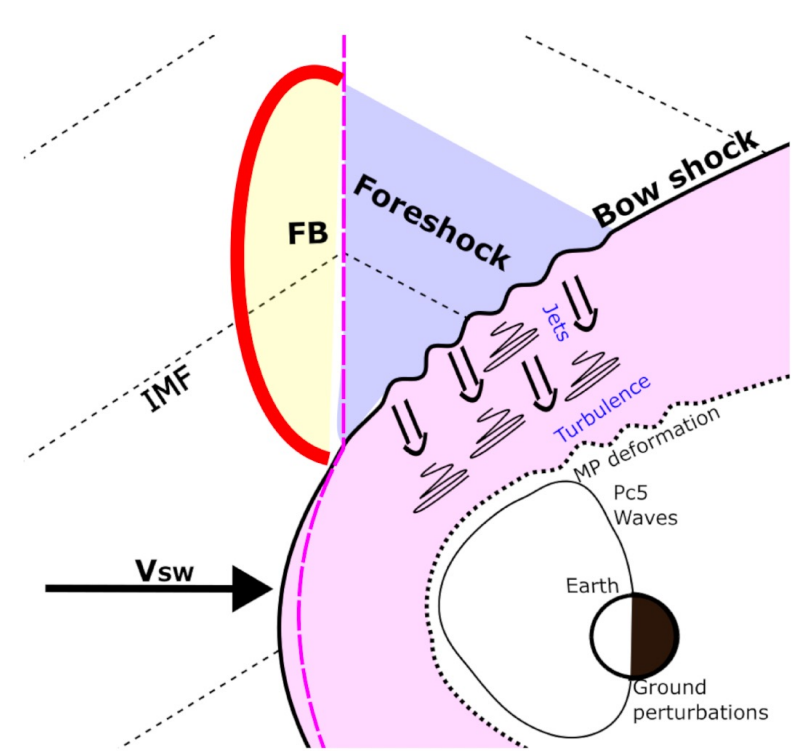
# Foreshock Transients



Travelling Foreshocks and Foreshock Compressive Boundaries

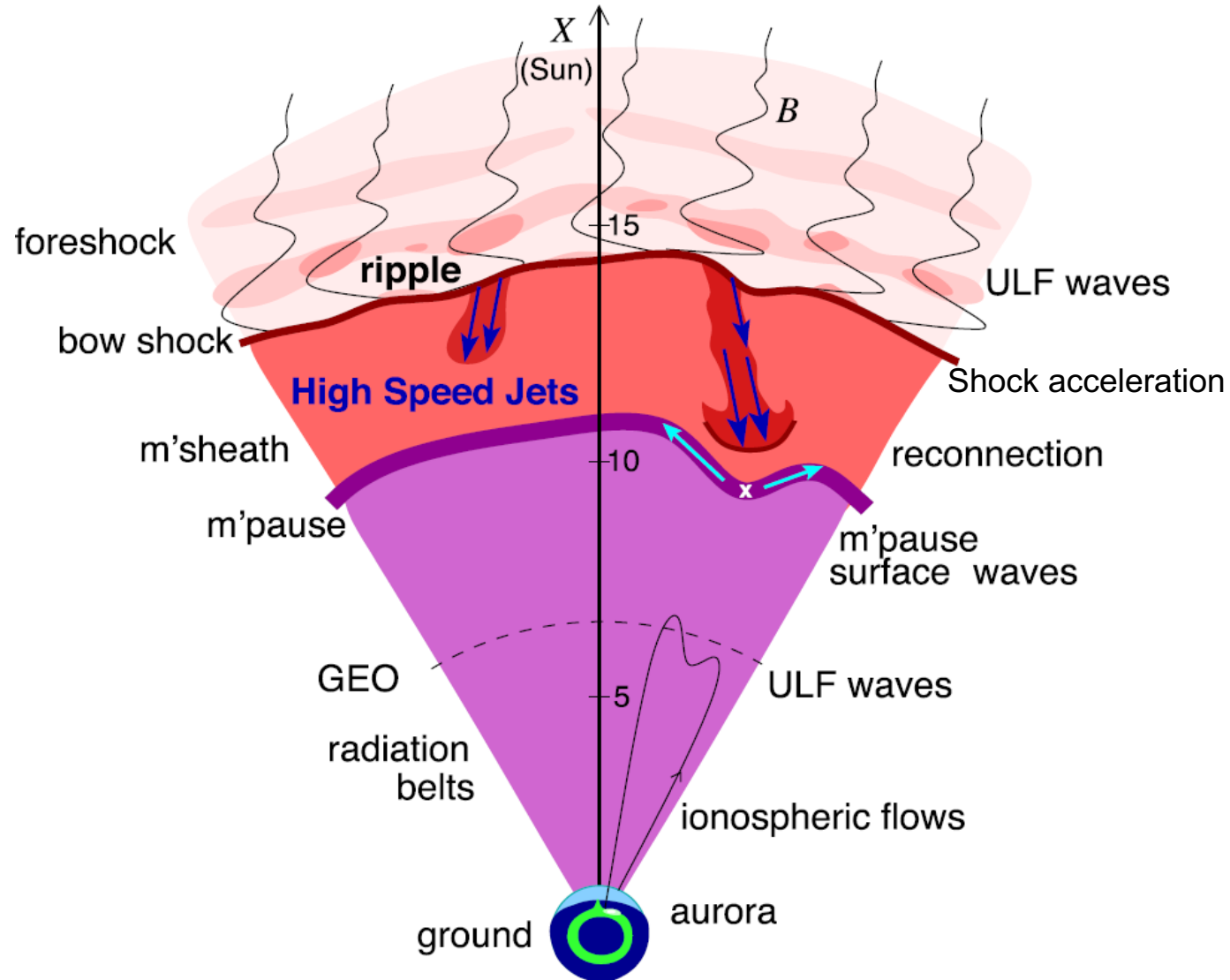


Hot Flow Anomalies



Foreshock Bubbles

# Magnetosheath jets



## Definition

Magnetosheath jets are **transient localized enhancements of dynamic pressure** (density and/or velocity increase)

*e.g., 200% dynamic pressure enhancement compared to background magnetosheath*

## Related phenomena

*Radiation belts  
Throat aurora  
Magnetopause reconnection  
Magnetopause penetration  
Shock acceleration  
Magnetopause surface eigenmodes  
ULF waves  
Substorms  
Ground magnetometer detection*

# Jets – references update (>2019)

## Associated phenomena & effects

- **Excitation** of surface **eigenmodes** at magnetopause: Archer et al. (2019, 2021)
- **Mirror mode waves** and jets: Bianco-Cano et al. (2020)
- Bursty **magnetic reconnection** at the Earth's magnetopause: Ng et al. (2021)
- **Ground-based magnetometer** response: Norenus et al. (2021)
- Generation of **Pi2 pulsations**: Katsavrias et al. (2021)
- B in jets, **Bz variations near magnetopause**: Vuorinen et al. (2021)
- High-Speed Jets **Triggering Dayside Ground ULF**: Wang et al. (2022)
- **Waves** and **jets** using burst MMS data: Krämer et al. (2023)

~110 citations

Jets Downstream of Collisionless Shocks

Plaschke et al. (2018)

<https://link.springer.com/article/10.1007/s11214-018-0516-3>

## Modeling & formation

- **Velocity & magnetic field alignment** in jets: Plaschke et al. (2020)
- **Classification** of jets using MMS & Neural Networks: Raptis et al. (2020a,2020b)
- Comparison **MMS vs simulations**: Palmroth et al. (2021)
- **Solar wind effect** on jet formation: LaMoury et al. (2021)
- Magnetosheath Jets and **Plasmoids** - Hybrid Simulations: Preisser et al. (2020)
- **Formation** of jets in **Quasi-perpendicular magnetosheath**: Primoz et al. (2021)
- **Occurrence** in relation to **CMEs and SIRs**: Koller et al. (2022)
- **Shock reformation** and the **formation of high-speed jets**: Raptis et al. (2022a)
- **Electron acceleration** and **bow waves** in jets: Vuorinen et al. (2022)
- **Kinetic structure** of jets and **partial plasma moments**: Raptis et al. (2022b)

And more : Liu et al. (2020a,2020b), Omelchenko et al (2021), Sibeck et al. (2021), Suni et al. (2021), Tinoco-Arenas et al. (2022) ... etc.

# Jets – references update (>2019)

~110 citations

Jets Downstream of Collisionless Shocks

Plaschke et al. (2018)

<https://link.springer.com/article/10.1007/s11214-018-0516-3>



**Jets Downstream of Collisionless Shocks:  
The Last Five Years**

Ongoing review (TBD)

# Dayside transient community & expectations from MMS

1. ISSI Science International Team 555 “**Impact of Upstream Mesoscale Transients on the Near-Earth Environment**”
2. The **Graz Jet Workshop** (Group working on magnetosheath jets)
3. ISSI Science International Team 546 “**Magnetohydrodynamic Surface Waves at Earth's Magnetosphere (and Beyond)**”

Martin Archer (leader)	Imperial College London, UK	3
Katariina Nykyri (co-leader)	Embry-Riddle Aeronautical University, USA	
Simone Di Matteo	Catholic University of America, USA	
Tom Elsdén	St Andrews University, UK	
Megan Gillies	University of Alberta, USA	
Michael Hartinger	Space Science Institute, USA	
Anatoly Leonovich	Institute of Solar-Terrestrial Physics, Irkutsk, Russia	
Bo Li	Shandong University Weihai, China	
Valery Nakariakov	University of Warwick, UK	
Vyacheslav Pilipenko	Space Research Institute, Moscow, Russia	
Ferdinand Plaschke	Technische Universität Braunschweig, Germany	
Xueling Shi	Virginia Tech, USA	
Kareem Sorathia	Johns Hopkins University Applied Physics Laboratory, USA	
Maria Walach	Lancaster University, UK	
Zhonghua Yao	Chinese Academy of Sciences, Beijing, China	
Kevin Blasl	Austrian Academy of Sciences, Graz, Austria	
Rachel Rice	University of Maryland, USA	
Frances Staples	University of California, Los Angeles, USA	

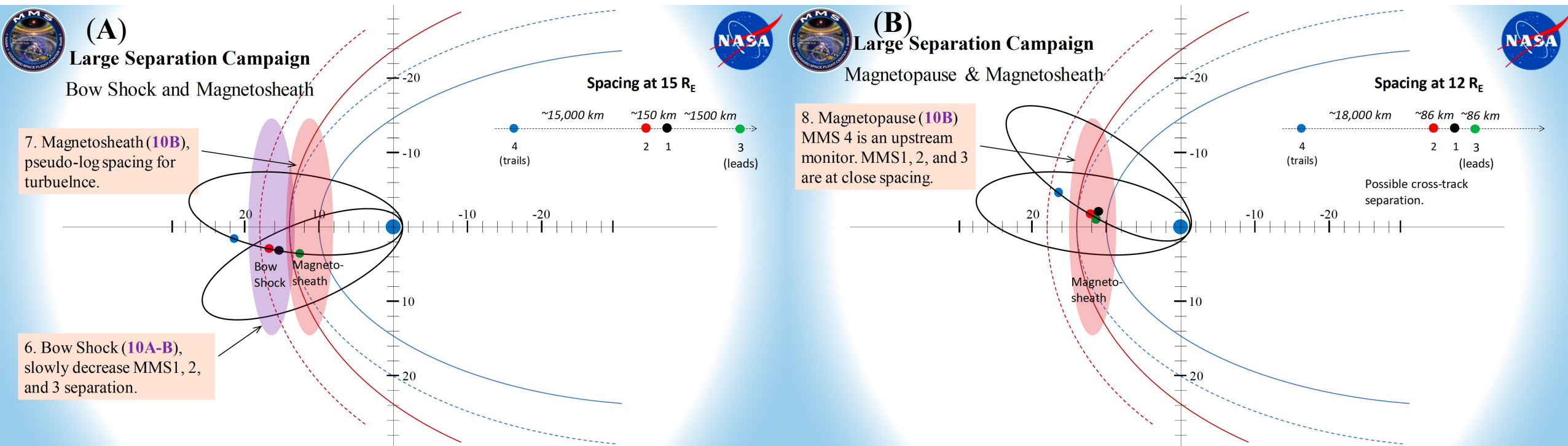
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Xóchitl Blanco-Cano	Geophysics Institute, UNAM, Mexico	
Lucile Ture	University of Helsinki, Finland	
Yann Pfau-Kempf	University of Helsinki, Finland	
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Adrian LaMoury	Imperial College London, UK	
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Marcos Vinicius Dias Silveira	University of Sao Paulo, EEL, Brasil	
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Sun Hee Lee	CUA, NASA GSFC, USA	

Savvas Raptis	ESA/ESTEC, Noordwijk, The Netherlands	2
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Laura Vuorinen	University of Turku, Finland	
Herbert Gunell	Umeå University, Sweden	
Eva Krämer	Umeå University, Sweden	
Niki Xirogiannopoulou	Charles University, Czech Republic	
Jonas Suni	University of Helsinki, Finland	
Adrian Pöppelwerth	TU Braunschweig, Germany	
Cyril Simon Wedlund	Space Research Institute, Austrian Academy of Sciences, Graz, Austria	
Manuela Temmer	Institute of Physics, University of Graz, Austria	
Luis Preisser	Space Research Institute, Austrian Academy of Sciences, Graz, Austria	
Tomas Karlsson	KTH Royal Institute of Technology, Sweden	
Florian Koller	Institute of Physics, University of Graz, Austria	
Xochitl Blanco-Cano	Instituto de Geofísica, UNAM, México	
Heli Hietala	Queen Mary University of London	

# What we wished for?

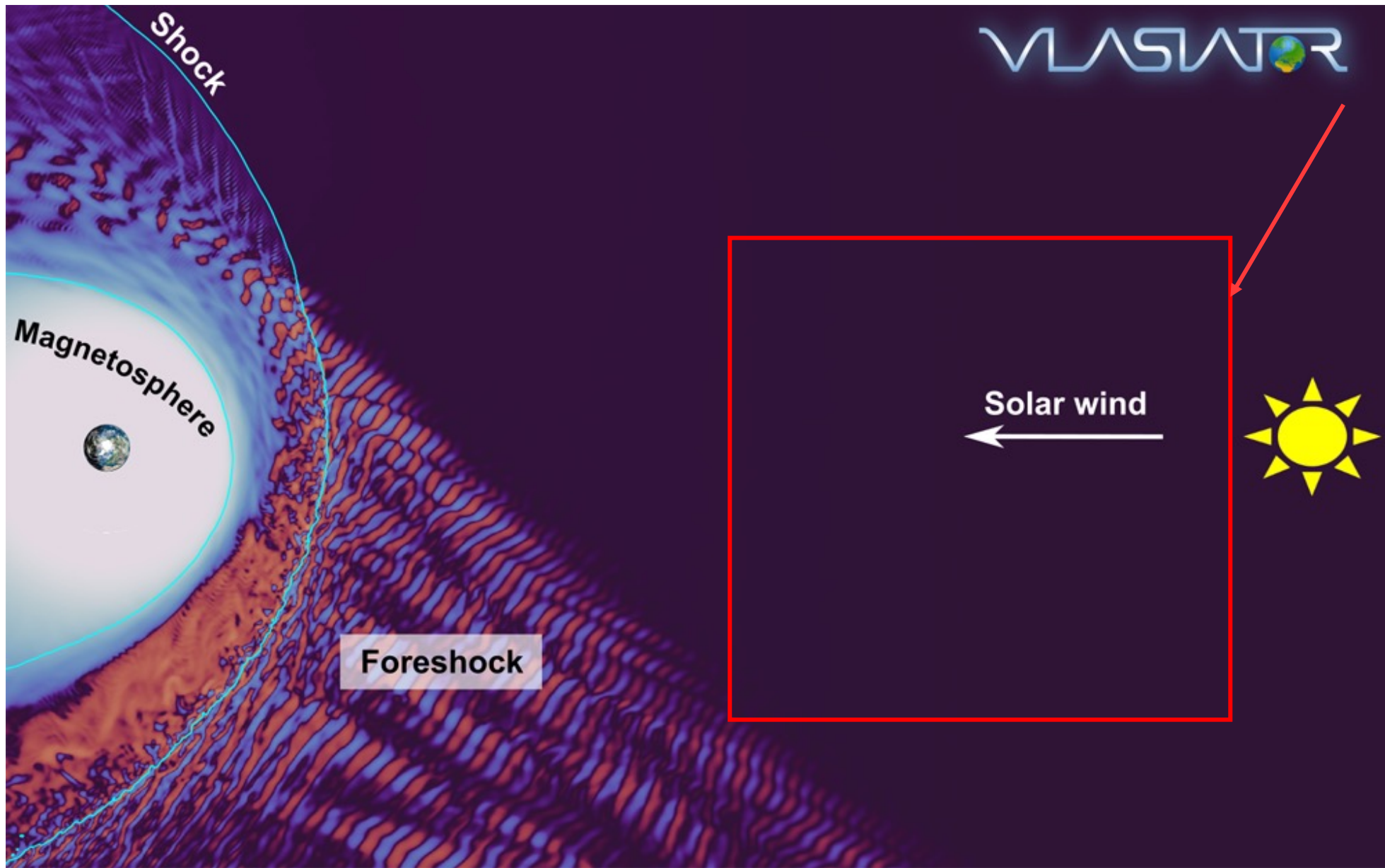
String-of-pearls campaigns with variable separation:

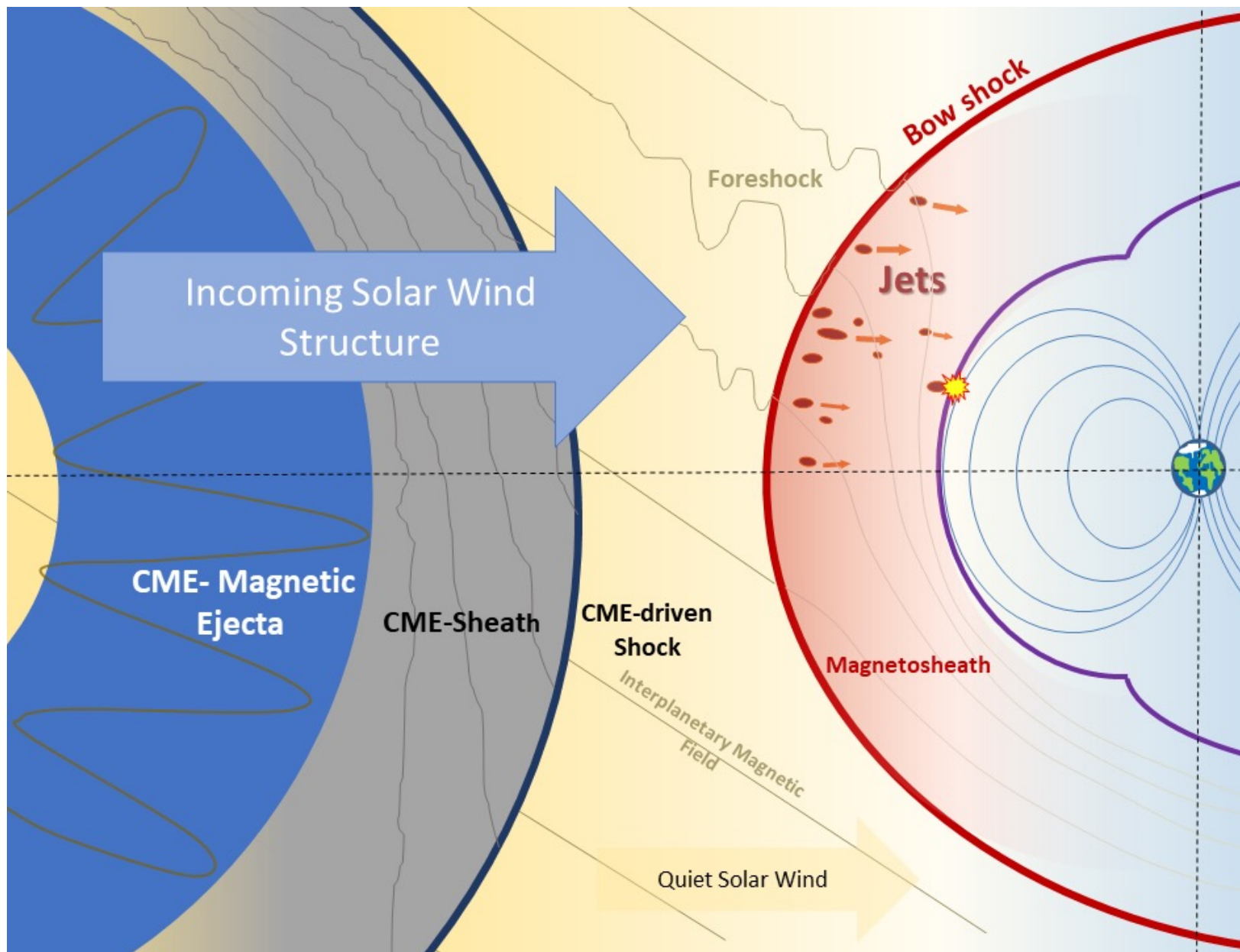
- (a) 1000s kms to investigate the effect of dayside transients on the magnetosphere
- (b) 1-2 satellites upstream and 1-2 downstream to evaluate the transmission and potential effects for the shock
- (c) Magnetopause response (1000s km separation) while upstream SW/foreshock (several  $R_E$ )





Why we need that ?  
(i.e., test cases)

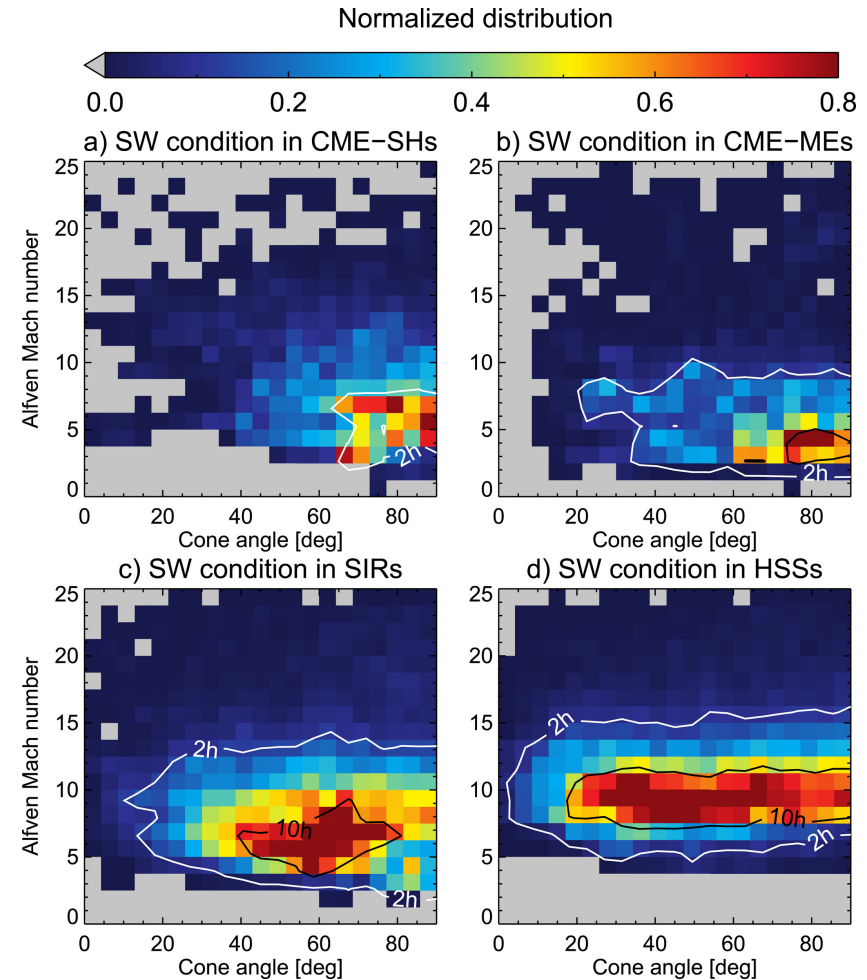
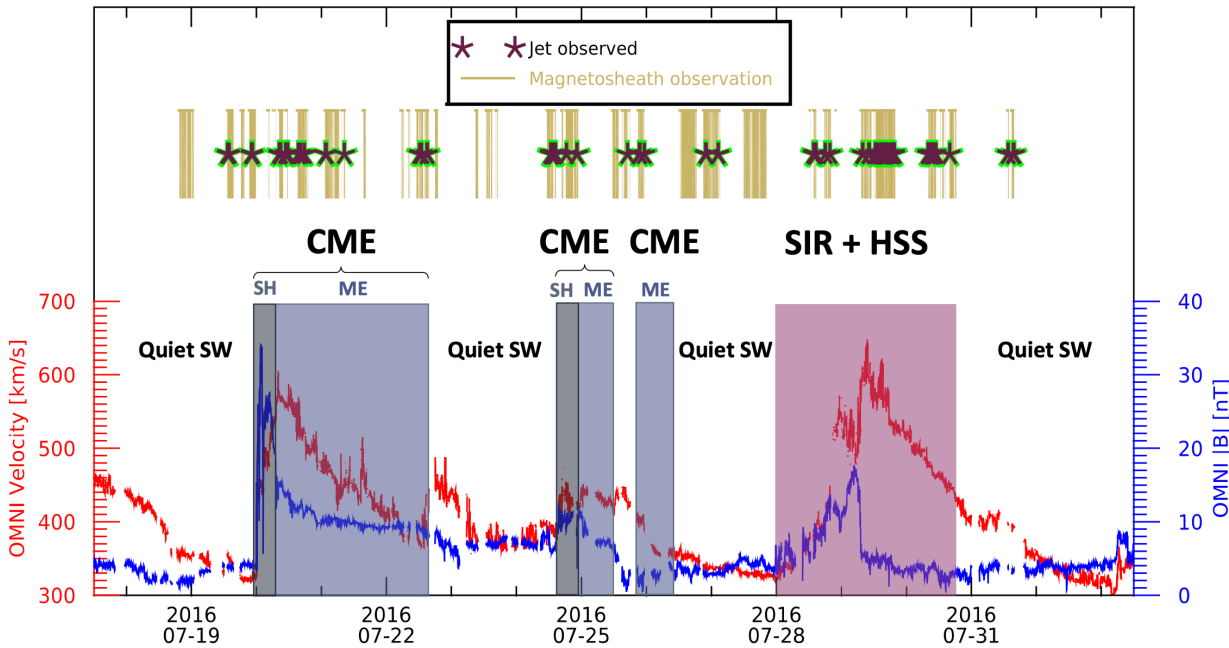




# Jet formation – Solar Transients

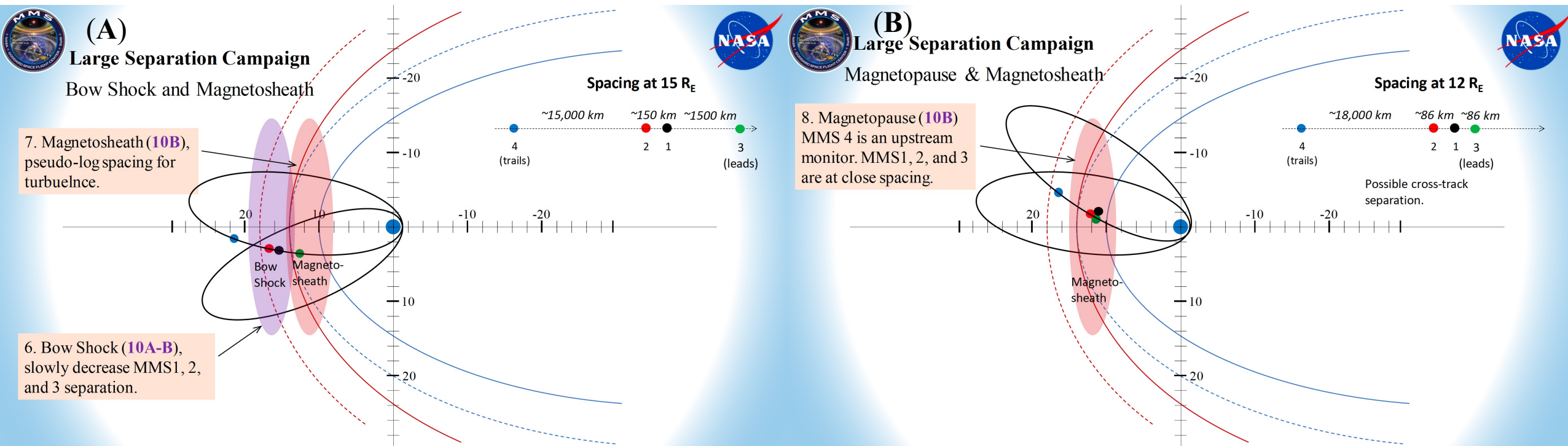
Magnetic Ejecta ↓ #jets  
 SIR + HSS ↑ #jets

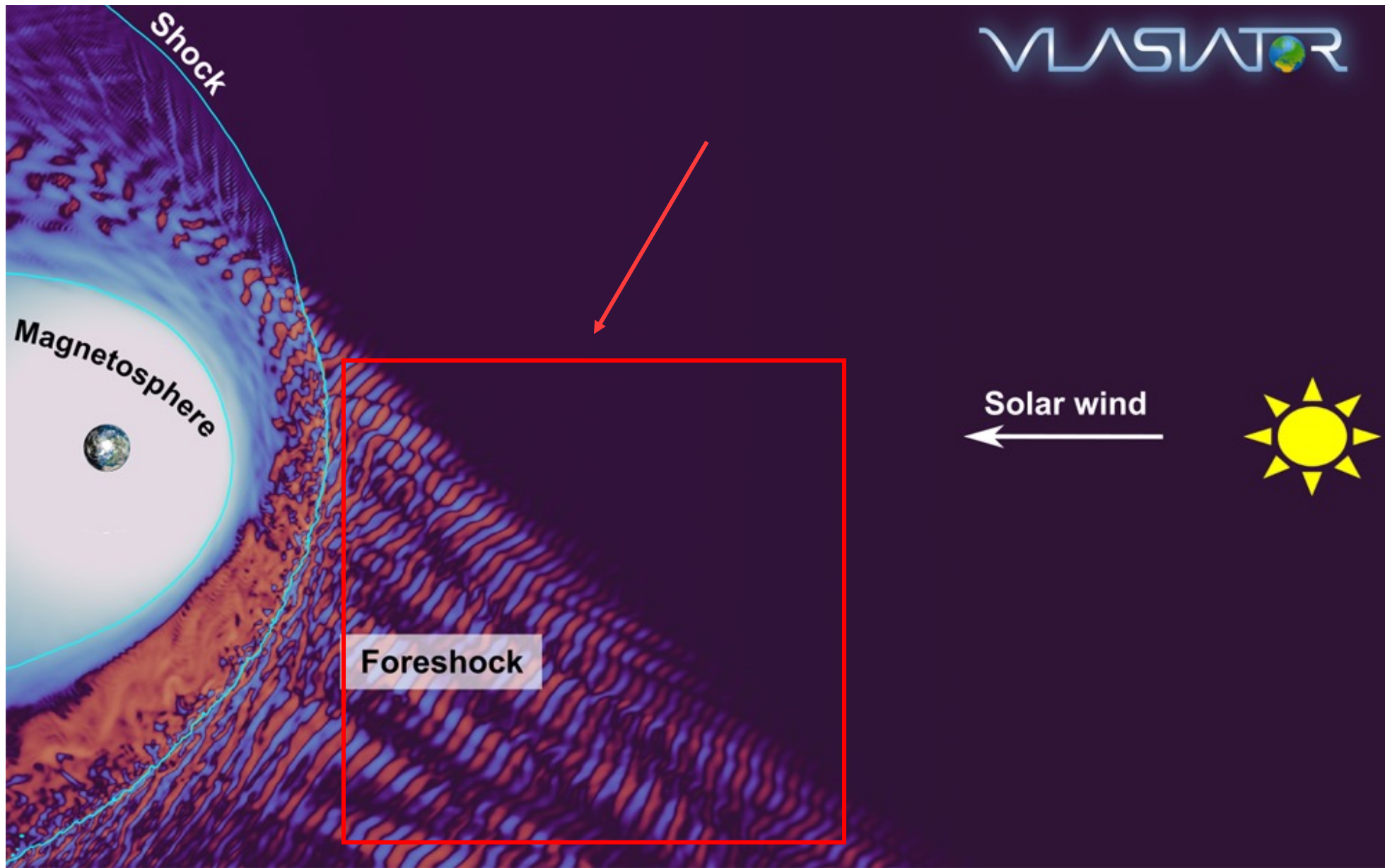
*Variable or extreme parameters  
 modify the occurrence and  
 properties of transients in the  
 foreshock / magnetosheath*



# Open questions on Solar transients

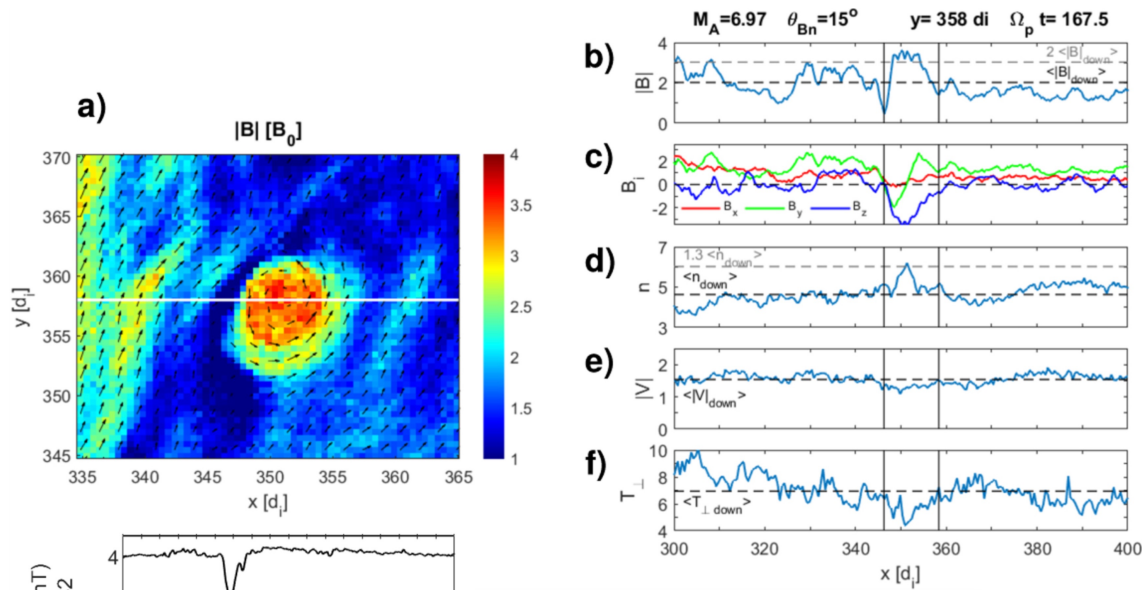
- 1) How do close to shock conditions vary with the presence of SW transient phenomena (magnetic clouds, CMEs, high speed streamers etc.) ? **(A)**
- 2) How does the formation of foreshock and magnetosheath transient change when there are SW transients? **(A)**
- 3) What are the exact effects of the solar transients to the magnetopause **(B)**



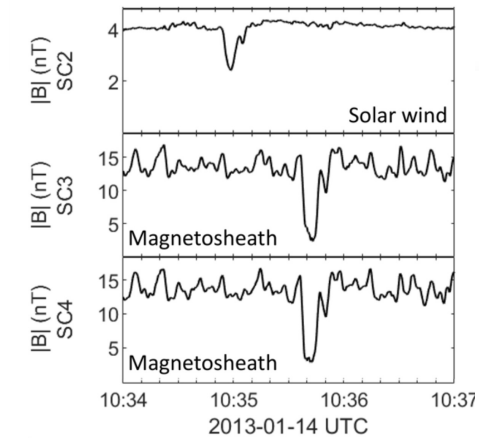


# Transmission of Foreshock Transients

ULF waves are transmitted and so are the non-linear associated phenomena (Shocklets, SLAMS, etc.)

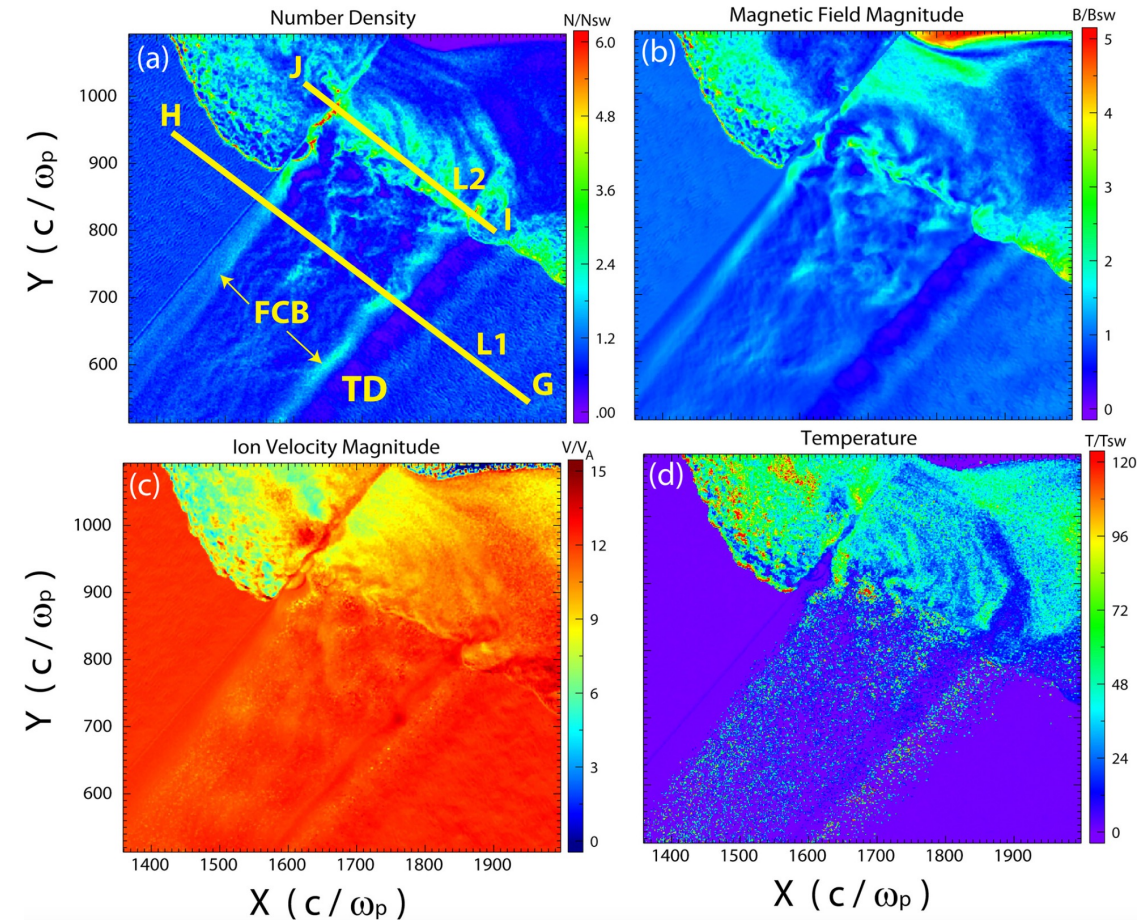


Preisser L., et al. 2020 | ApJL



Karlsson T., et al. 2022 | ANGE0

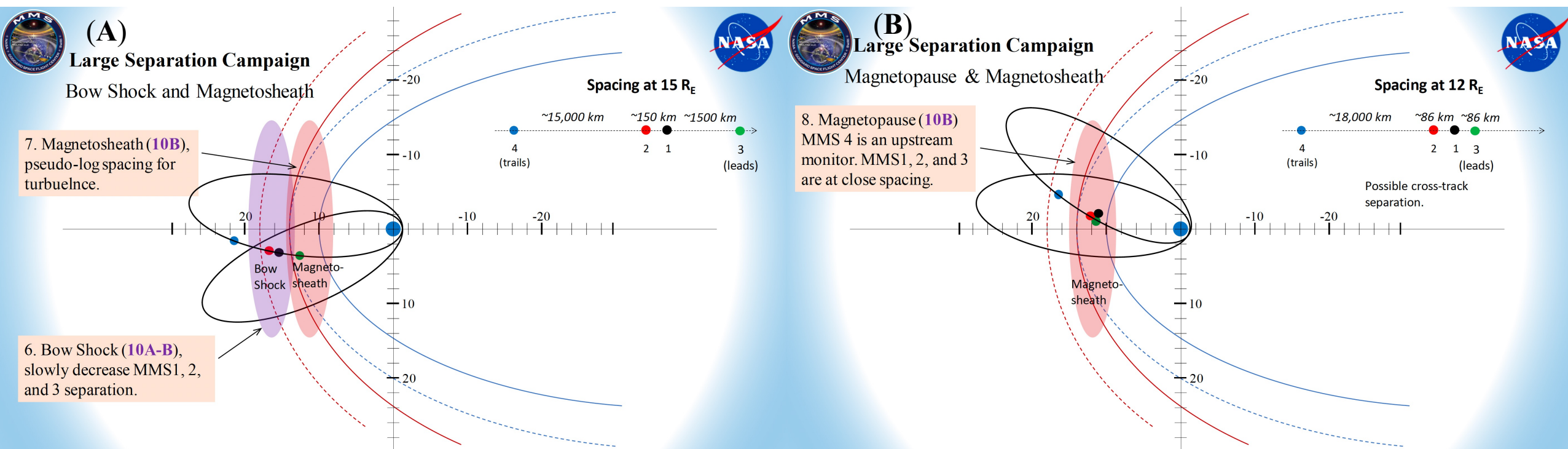
Transmission of FCB, FBs, HFAs, etc. has been shown in simulation and observations.



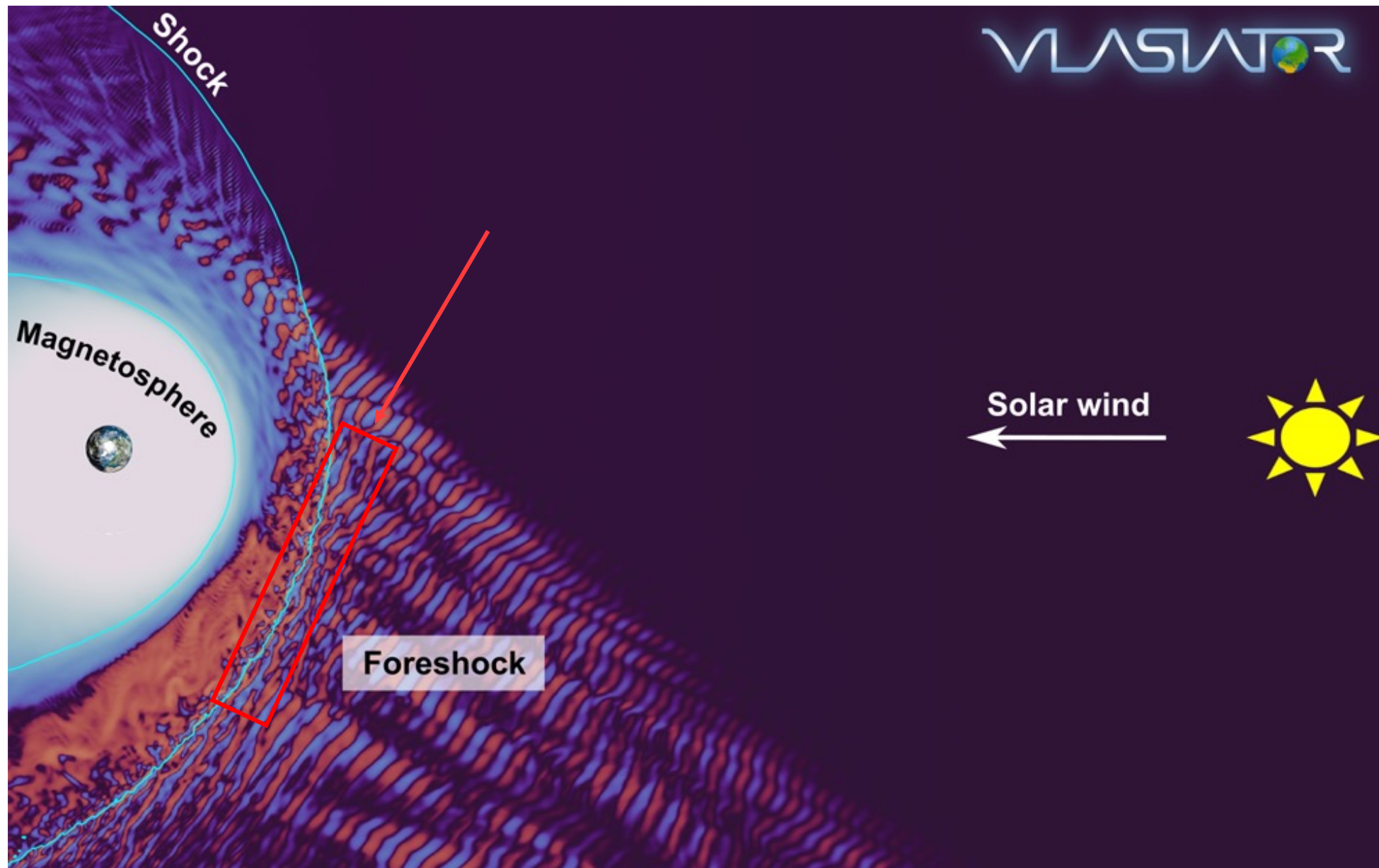
Sibeck D., et al. 2021 | JGR

# Open questions on Foreshock transients

1. What are the observational signatures of ULF (and associated phenomena) transmission ? How are the properties changing ? (A)
2. What effects can be caused in the magnetosheath plasma and in the magnetopause through the interaction of these structures with the ambient plasma (A, B)

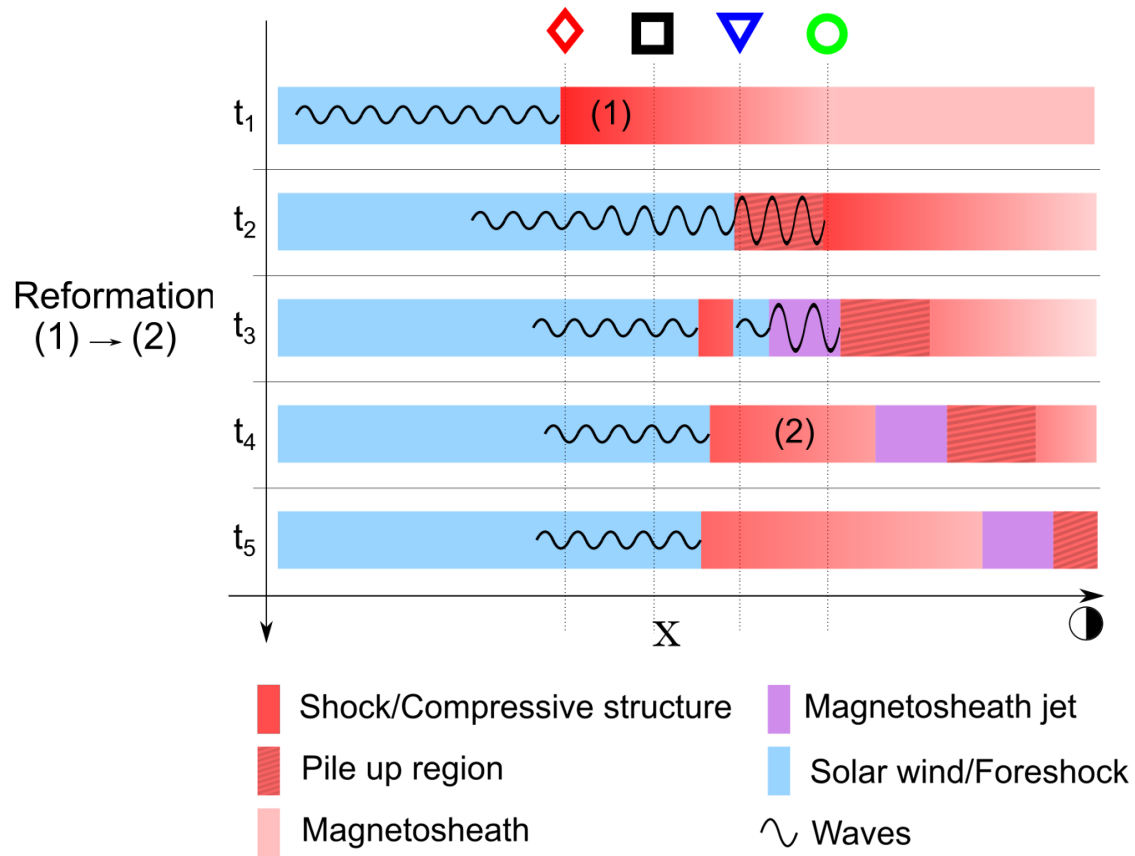






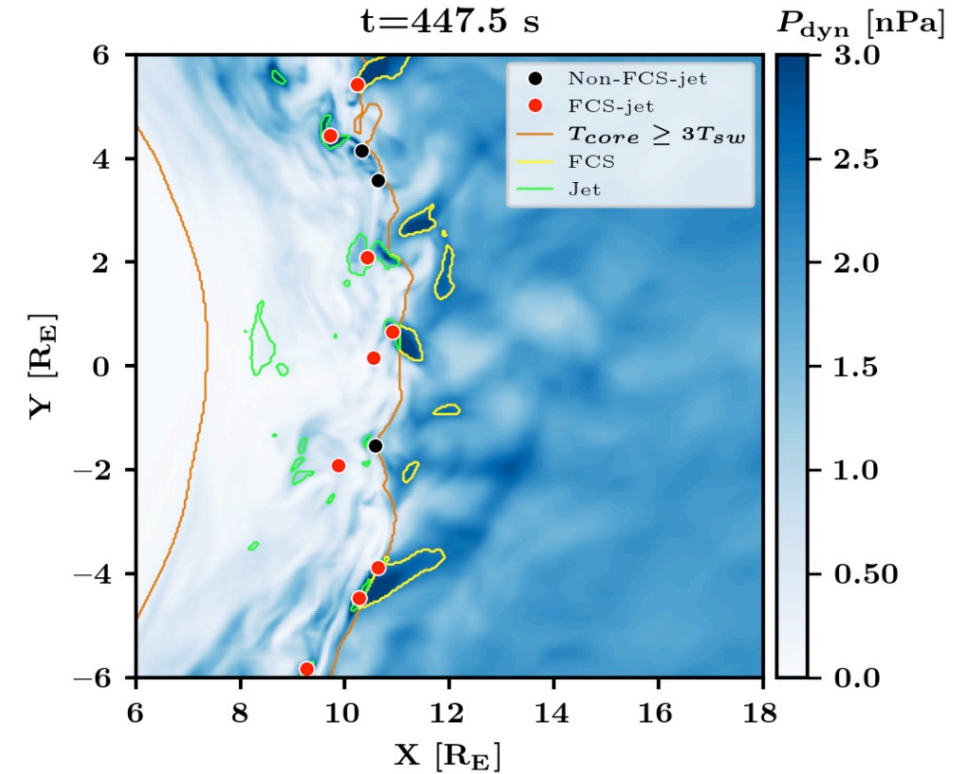
# Jet formation

*Jets forming from shock's fundamental non-stationarity*



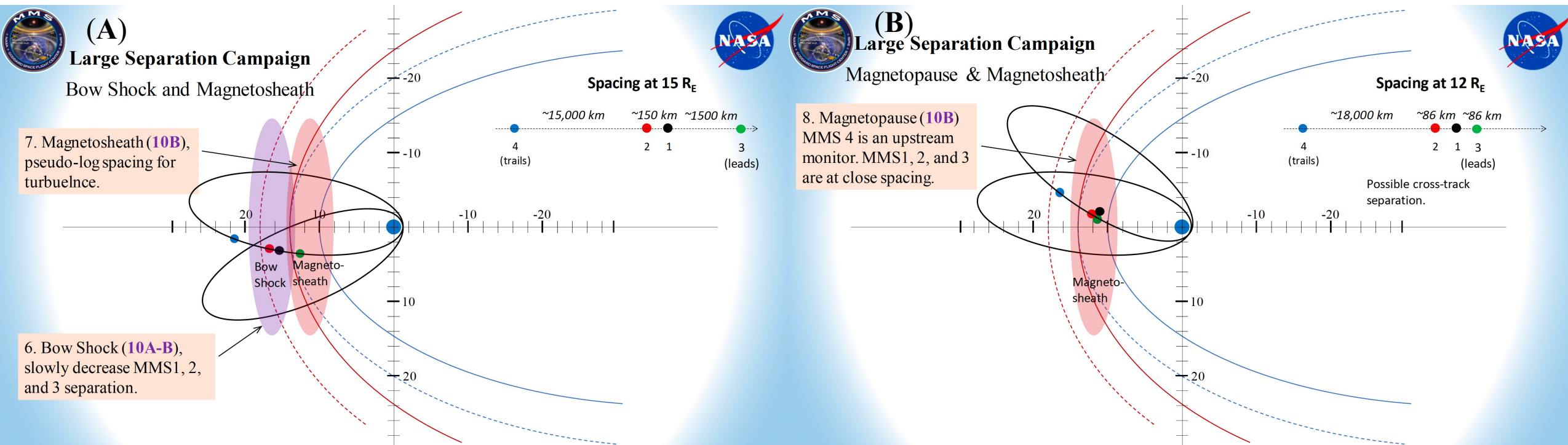
Turbulence Campaign 2019

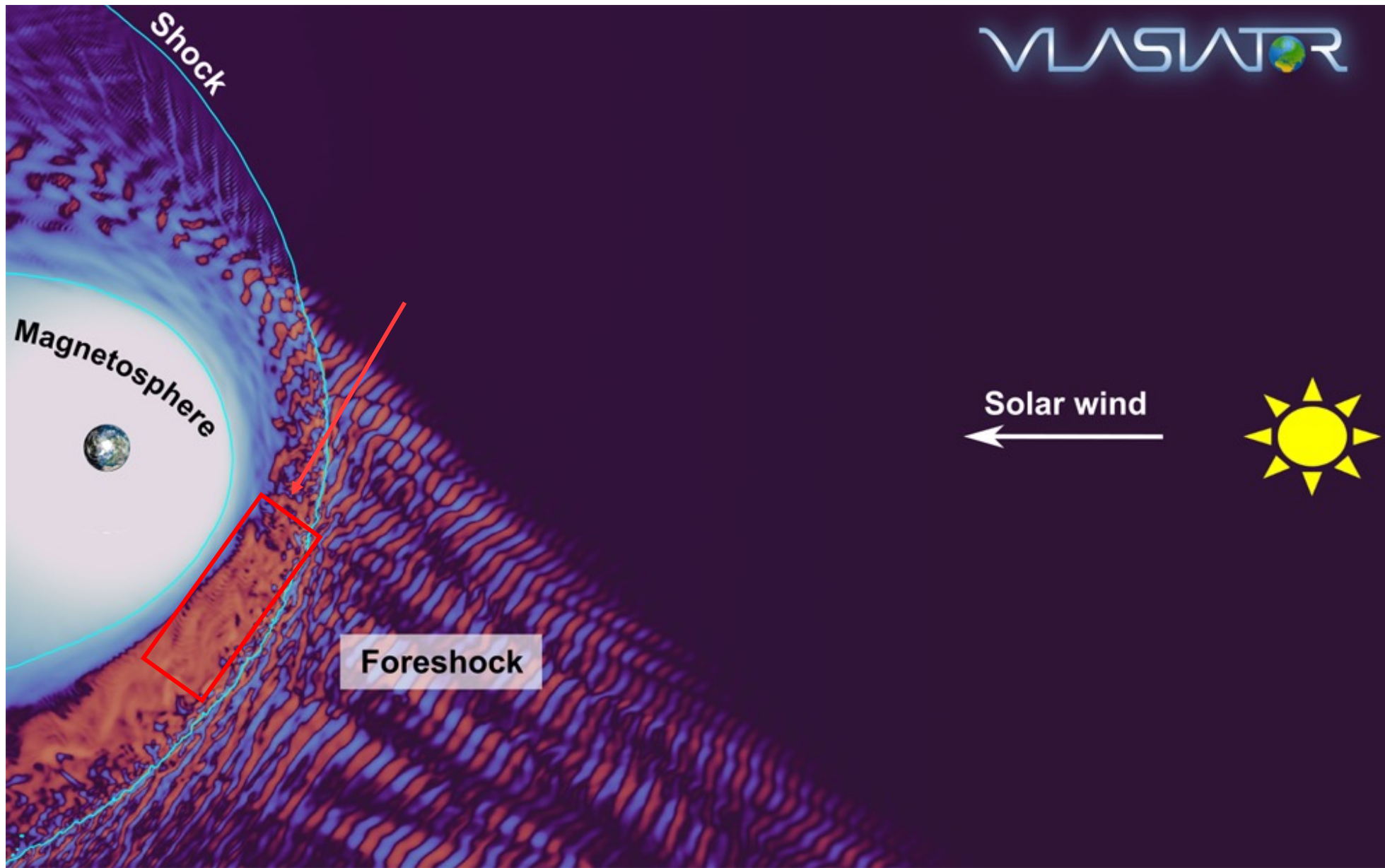
*75% of jets caused by Foreshock Compressive Structures*



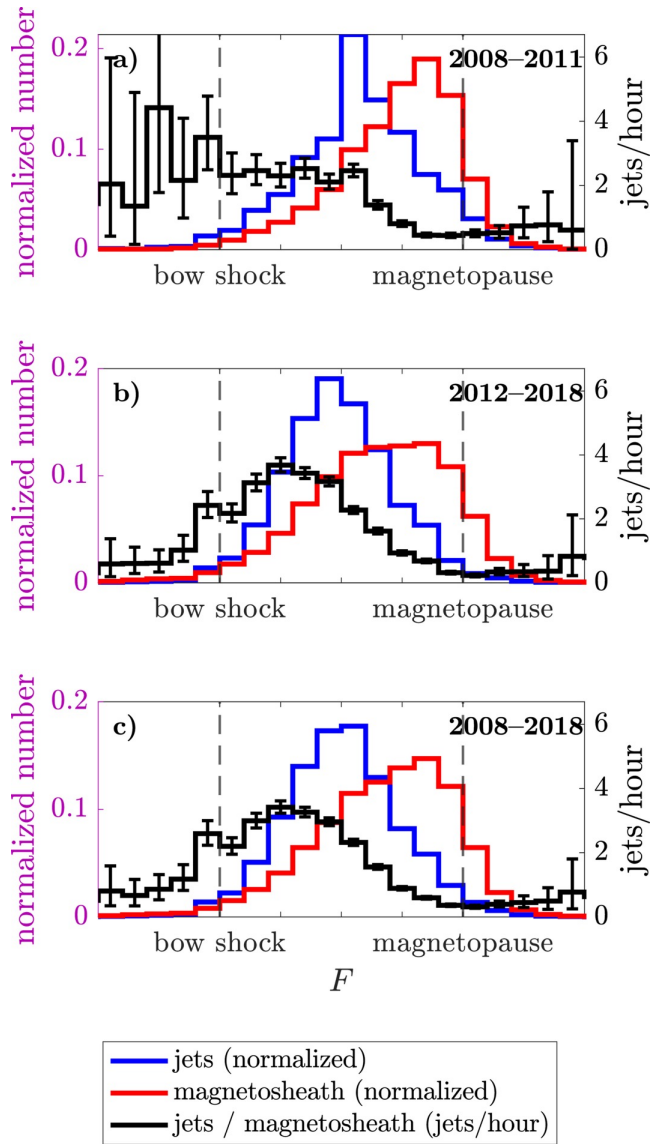
# Open questions on formation

- 1) Is global shock reformation the answer to jets ? or maybe faster flows get dissipated and become part of the background population before reaching the MP ? (A)
- 2) How are close to the shock upstream conditions affecting the formation of downstream jets ? (A, B)





# Jet evolution

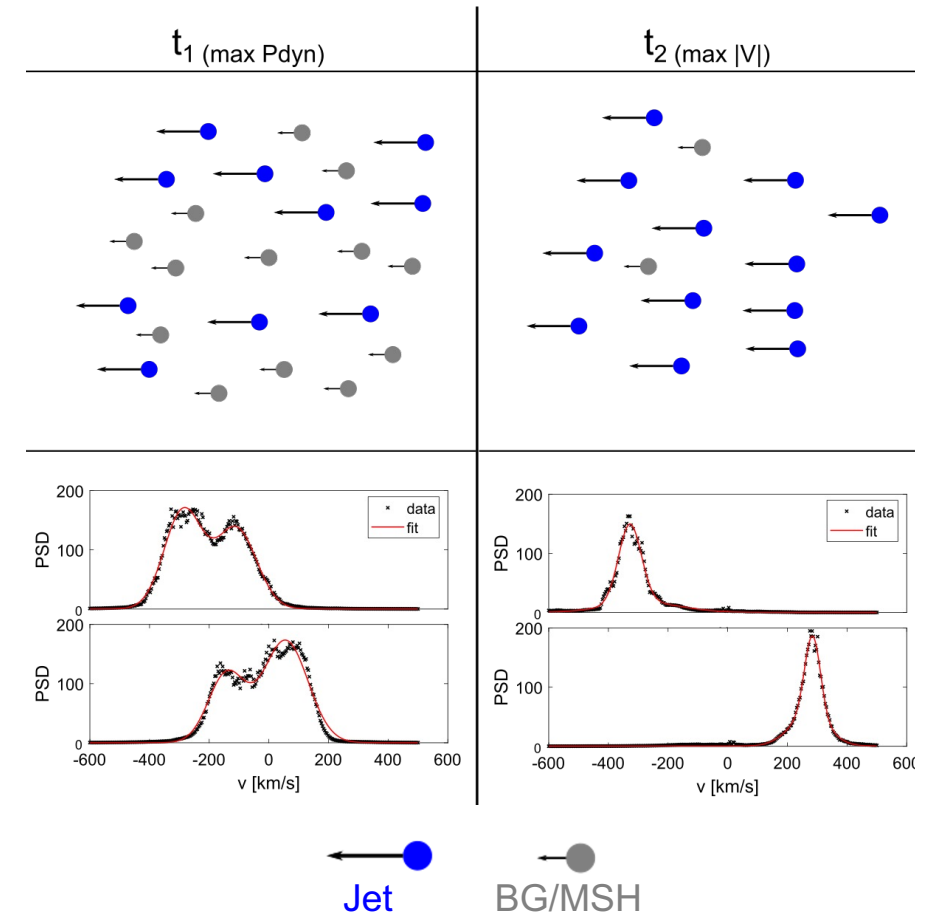


When can a jet reach the MP?

- 8 times more likely for high solar wind speed
- 17 more likely for low cone angles

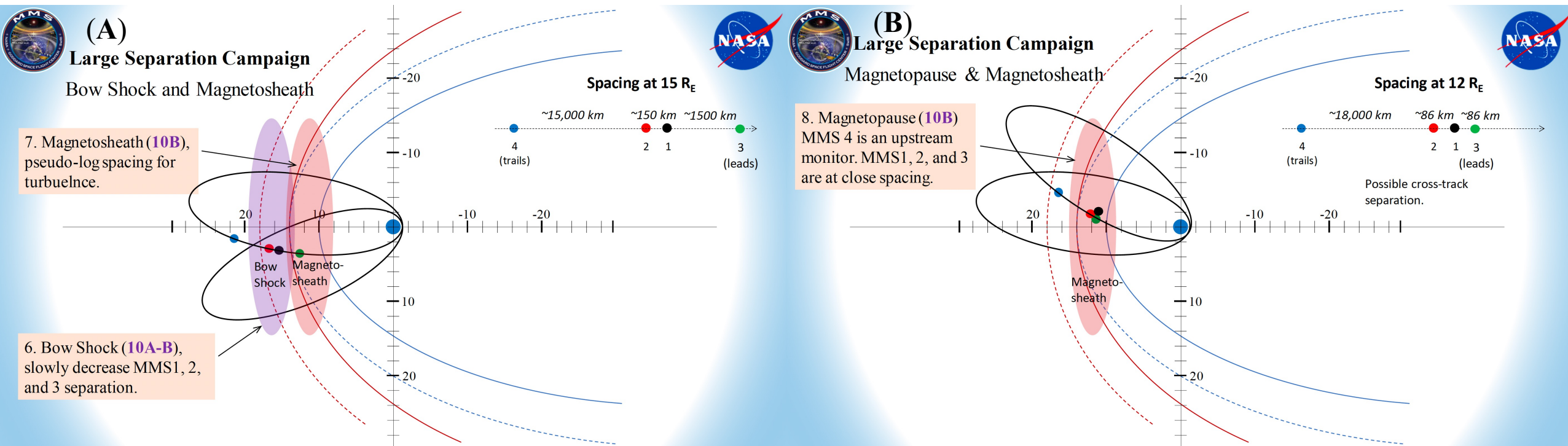
Jets **thermalize** towards the magnetopause, the shape of the jets **flatten** and they are able to **maintain their flow velocity and direction** within the magnetosheath flow

Jets exhibiting 2+ populations = partial moments needed = similar properties to upstream SW



# Open questions on Evolution

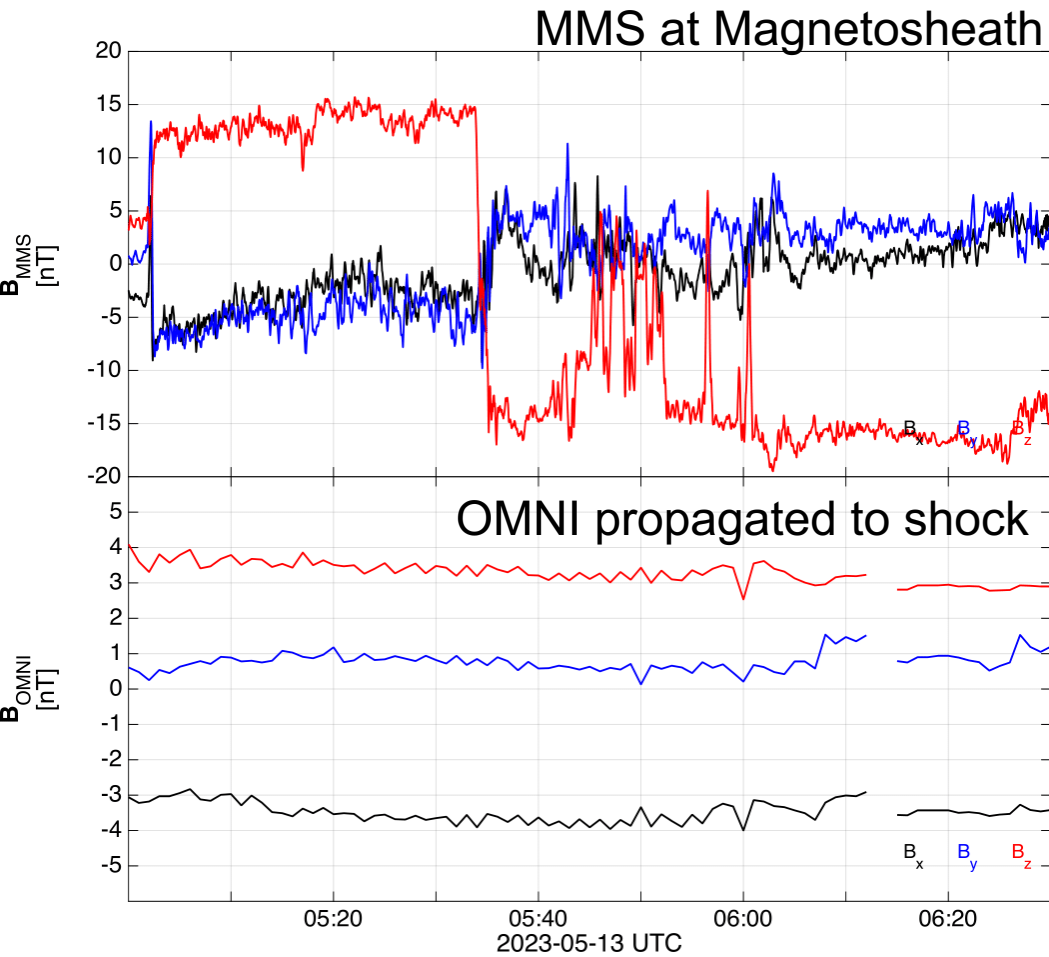
- 1) Do jets continue to show a clear double population? Does this change over time or closer to the MP ? (A)
- 2) How are the properties and the shape vary over distance from the bow shock/magnetopause? (A)
- 3) How do jets contribute to the generation of wave activity, turbulence and particle energization? (A)



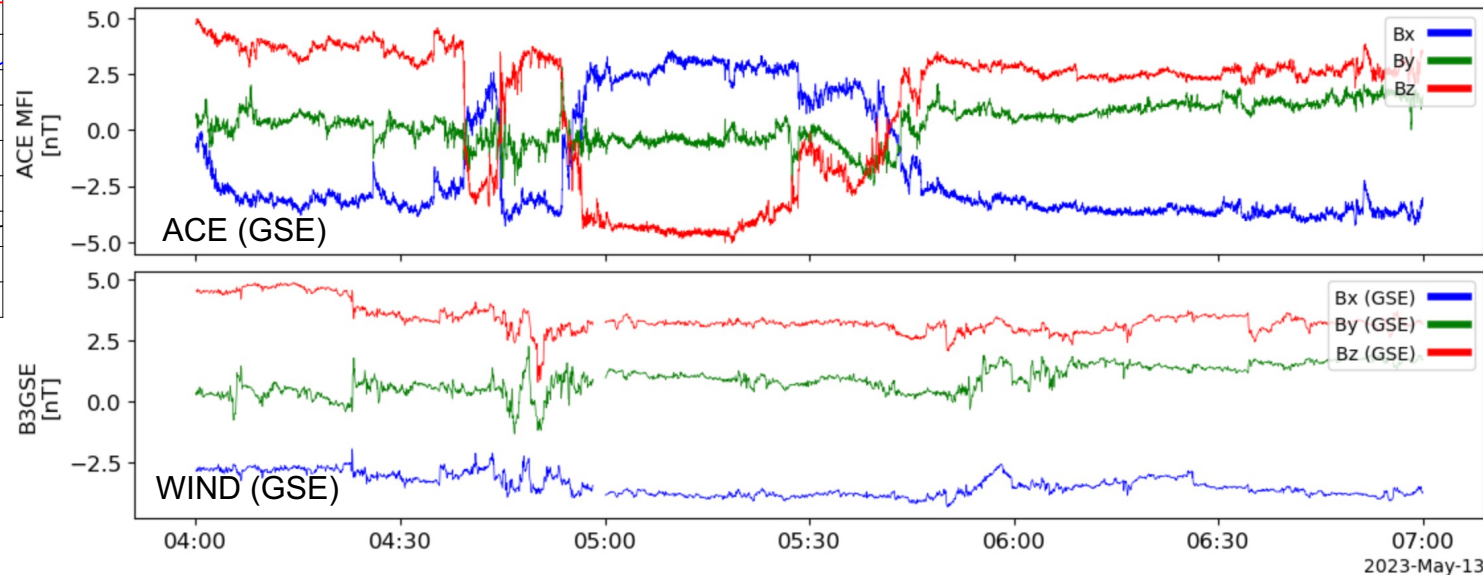
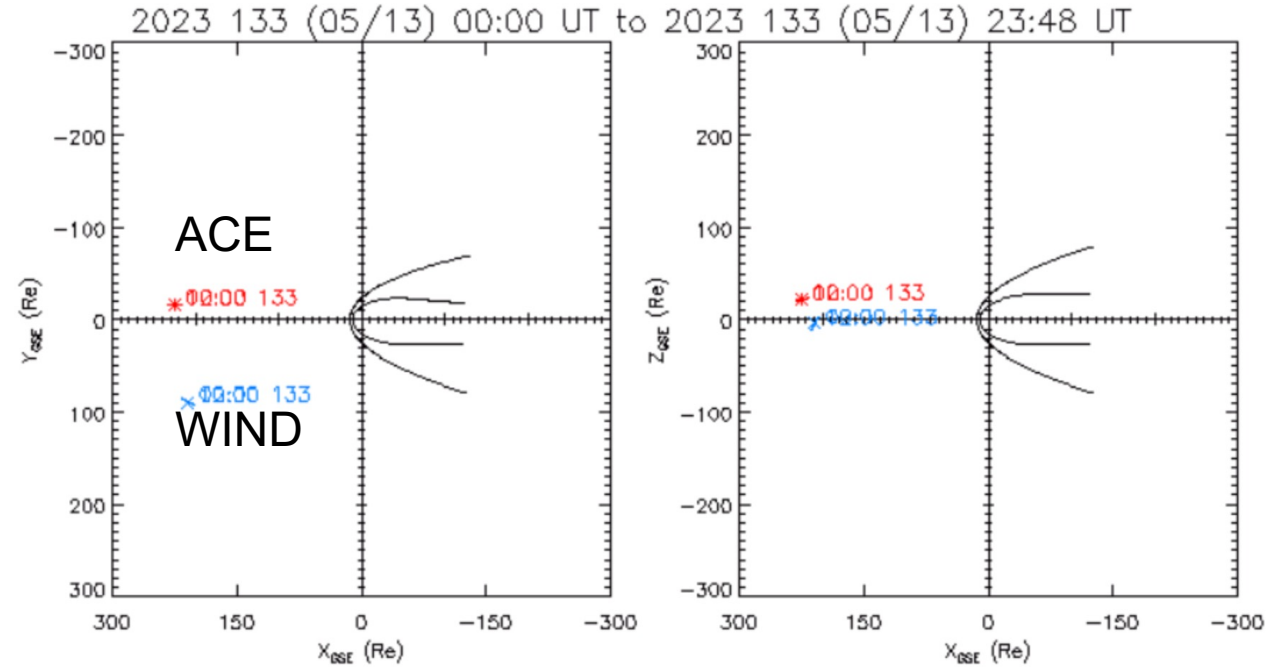
# Discussion

1. Foreshock & magnetosheath phenomena, in particular transient events = Interesting multi-scale physics, space weather effects, and overall unique environments to expand our understanding and test our codes.
2. MMS future campaigns: Opportunity to investigate different scales in high time-resolution, revealing previously unknown relationships between phenomena.
3. (Skipped): Effects on inner magnetosphere, excitation of ULF waves, etc.
4. (Skipped): Potential conjunctions with THEMIS/Cluster and other missions for investigating even larger scale evolution.

# +1 bonus reason for the upstream monitor campaign



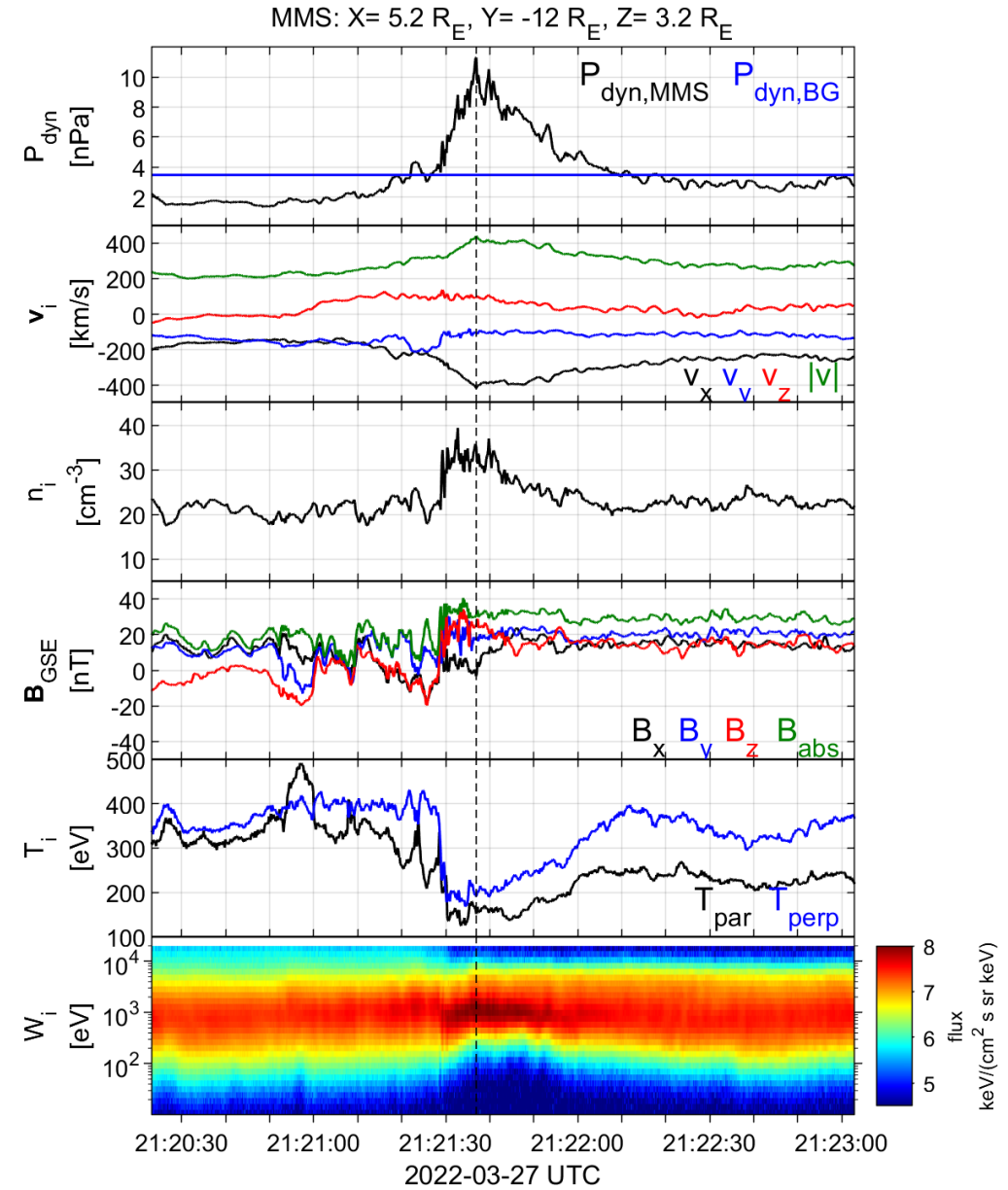
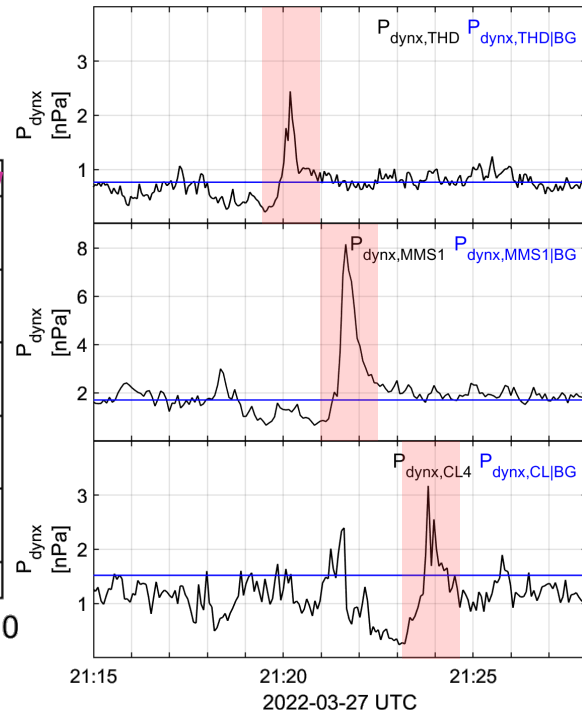
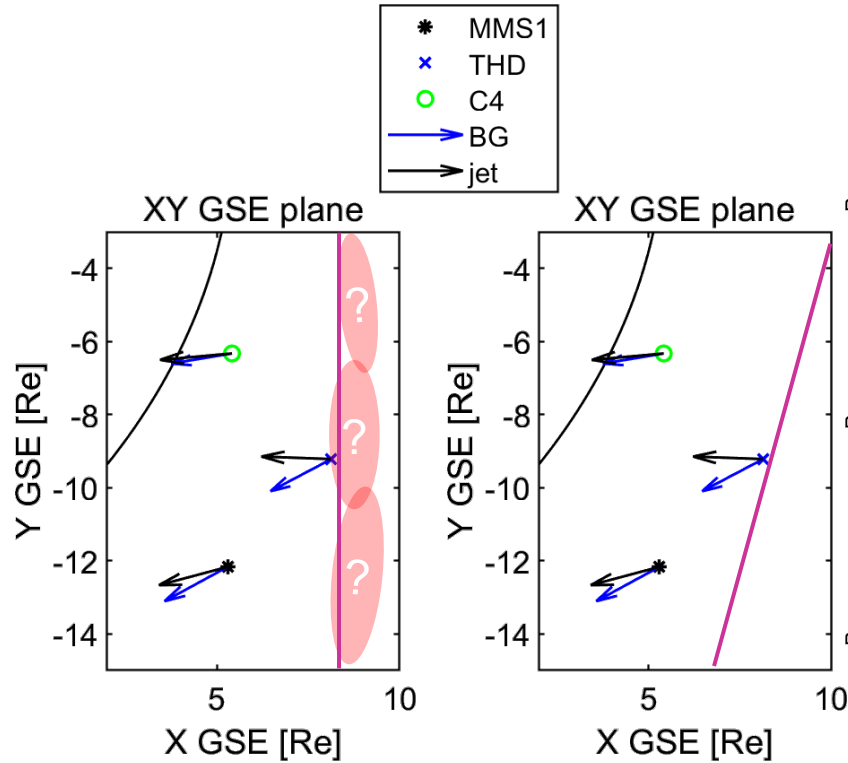
**Key-point:** Several minutes of Dt is typical, but cases like above are not rare either. Be careful when using propagated values





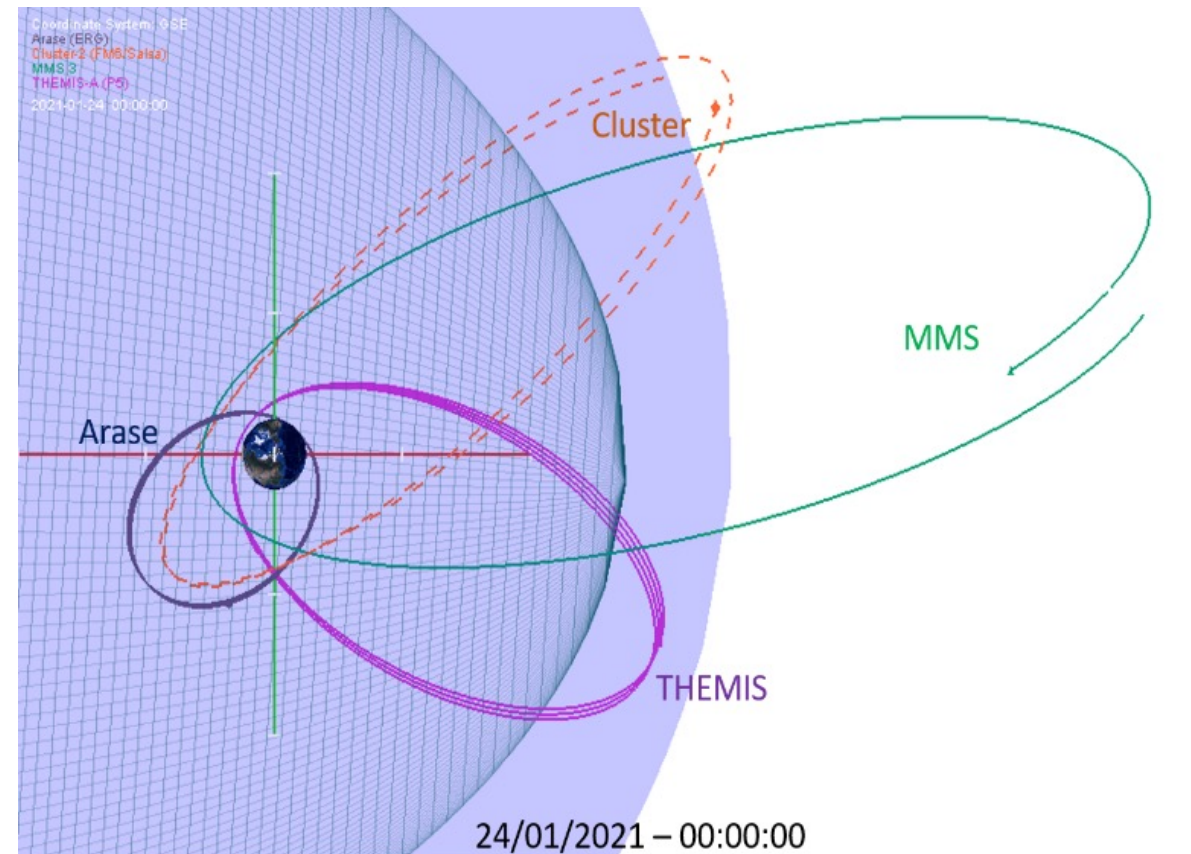
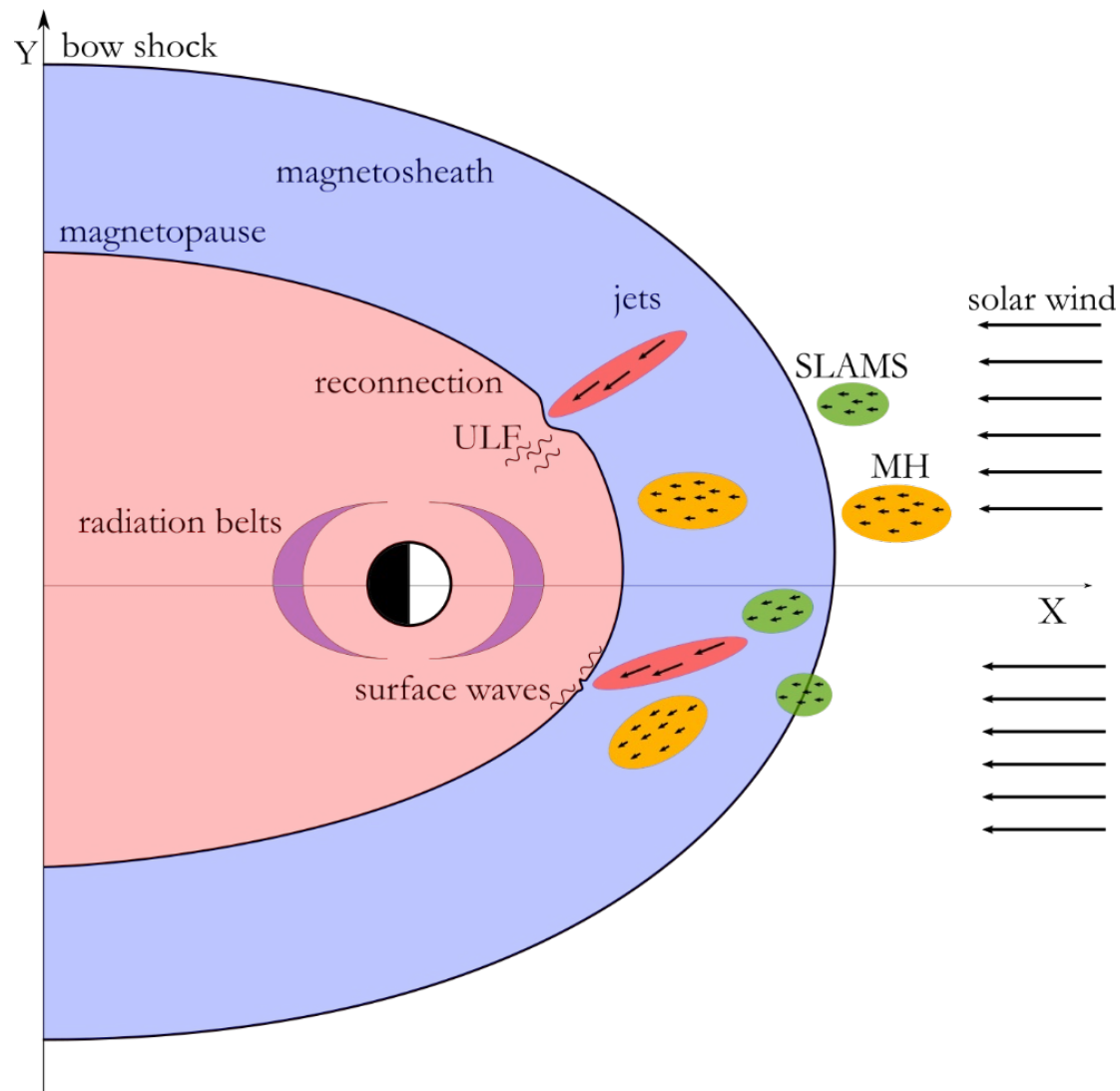
# Extras General

# Example of a multi-SC Jet

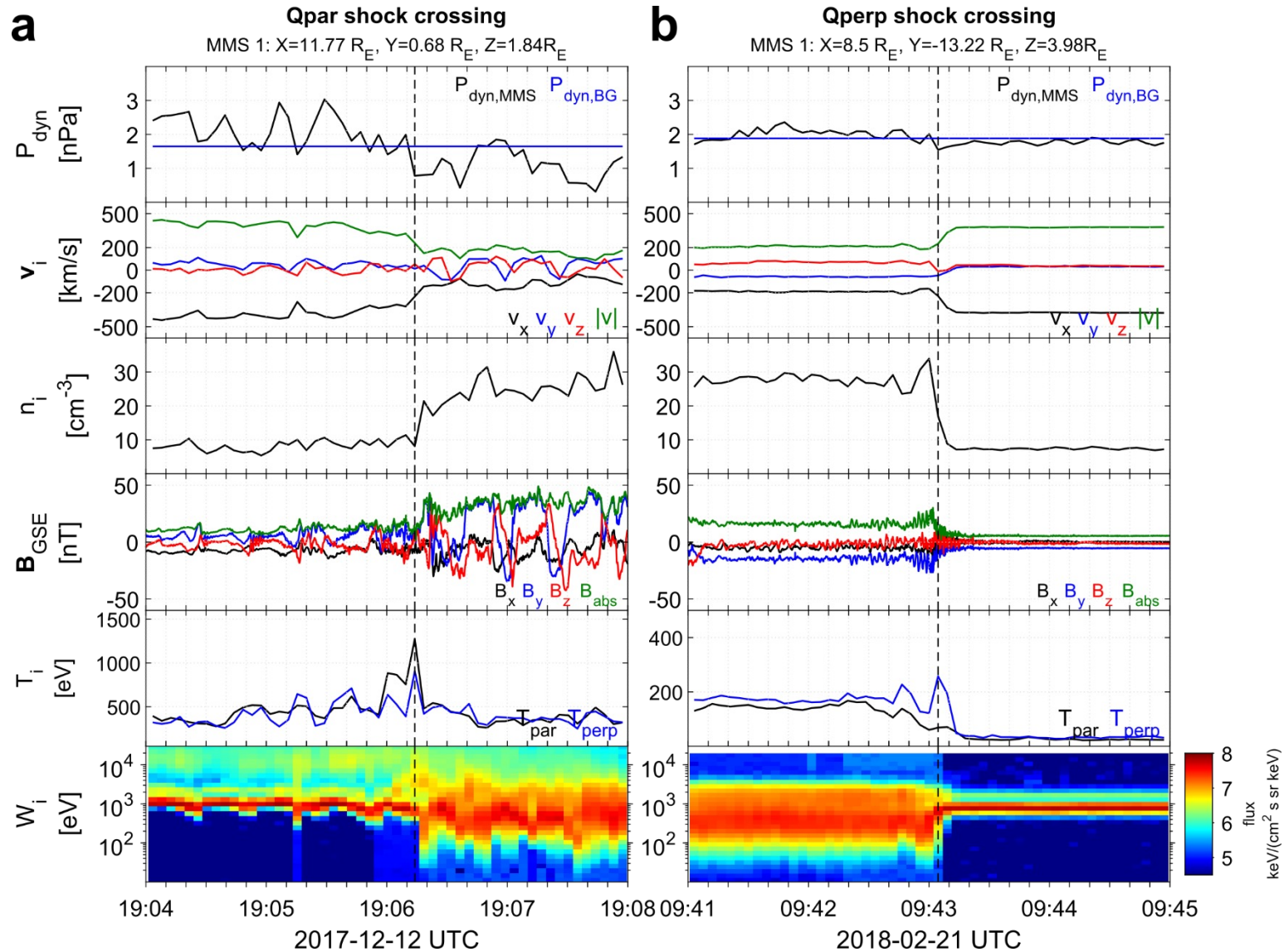


(Ongoing work – TBD)

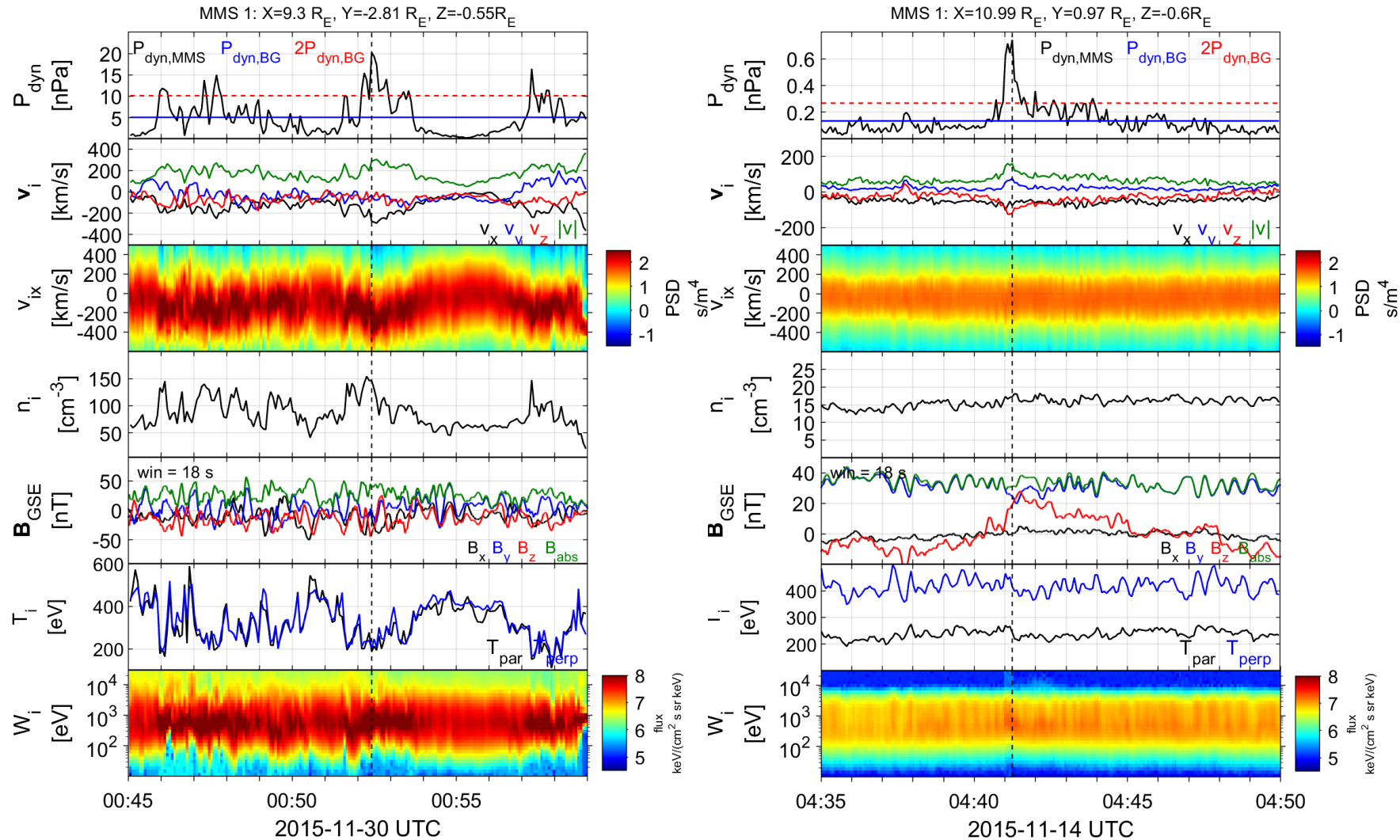
# A lot of data are not fully used (conjunctions)



# Qpar – Qperp crossings





# Qpar – Qperp Jets




# Why do we care? “big picture”

## Dayside Transient Phenomena and Their Impact on the Magnetosphere and Ionosphere

Hui Zhang , Qiugang Zong , Hyunju Connor, Peter Delamere, Gábor Facskó, Desheng Han, Hiroshi Hasegawa, Esa Kallio, Árpád Kis, Guan Le, Bertrand Lembège, Yu Lin, Terry Liu, Kjellmar Oksavik, Nojan Omidj, Antonius Otto, Jie Ren, Quanqi Shi, David Sibeck & Shutao Yao

*Space Science Reviews* **218**, Article number: 40 (2022) | [Cite this article](#)

## Transmission of foreshock waves through Earth’s bow shock

L. Turc , O. W. Roberts, D. Verscharen, A. P. Dimmock, P. Kajdič, M. Palmroth, Y. Pfau-Kempf, A. Johlander, M. Dubart, E. K. J. Kilpua, J. Soucek, K. Takahashi, N. Takahashi, M. Battarbee & U. Ganse

*Nature Physics* (2022) | [Cite this article](#)

## Downstream high-speed plasma jet generation as a direct consequence of shock reformation

Savvas Raptis , Tomas Karlsson, Andris Vaivads, Craig Pollock, Ferdinand Plaschke, Andreas Johlander, Henriette Trollvik & Per-Arne Lindqvist

*Nature Communications* **13**, Article number: 598 (2022) | [Cite this article](#)

EDI\*


Foreshock and magnetosheath transient phenomena and their effects on planetary magnetospheres.

Co-organized by PS2


Convener: Savvas Raptis  | Co-conveners: Heli Hietala , Ferdinand Plaschke , Tomas Karlsson , Christian Mazelle 

Abstract submission

## Geophysical Research Letters\*





Research Letter |  [Free Access](#)

### Investigating the Role of Magnetosheath High-Speed Jets in Triggering Dayside Ground Magnetic Ultra-Low Frequency Waves

Boyi Wang , Yukitoshi Nishimura, Heli Hietala, Vassilis Angelopoulos

First published: 07 November 2022 | <https://doi.org/10.1029/2022GL099768>

## Geophysical Research Letters\*

Research Letter |  [Open Access](#) |   

### Connection Between Foreshock Structures and the Generation of Magnetosheath Jets: Vlasiator Results

J. Suni , M. Palmroth, L. Turc, M. Battarbee, A. Johlander, V. Tarvus, M. Alho, M. Bussov, M. Dubart, U. Ganse, M. Grandin, K. Horaites, T. Manglayev, K. Papadakis, Y. Pfau-Kempf, H. Zhou



Impact of Upstream Mesoscale Transients on the Near-Earth Environment

ISSI team lead by Primož Kajdič & Xóchitl Blanco-Cano

## Foreshocks Across The Heliosphere: System Specific Or Universal Physical Processes?

ISSI Team led by H. Hietala (UK) & F. Plaschke (AT)



Co-organized by PS2

Convener: Savvas Raptis  | Co-conveners: Heli Hietala , Ferdinand Plaschke , Tomas Karlsson , Christian Mazelle 

Abstract submission