

# Shock Kinetic Processes and Particle Energization

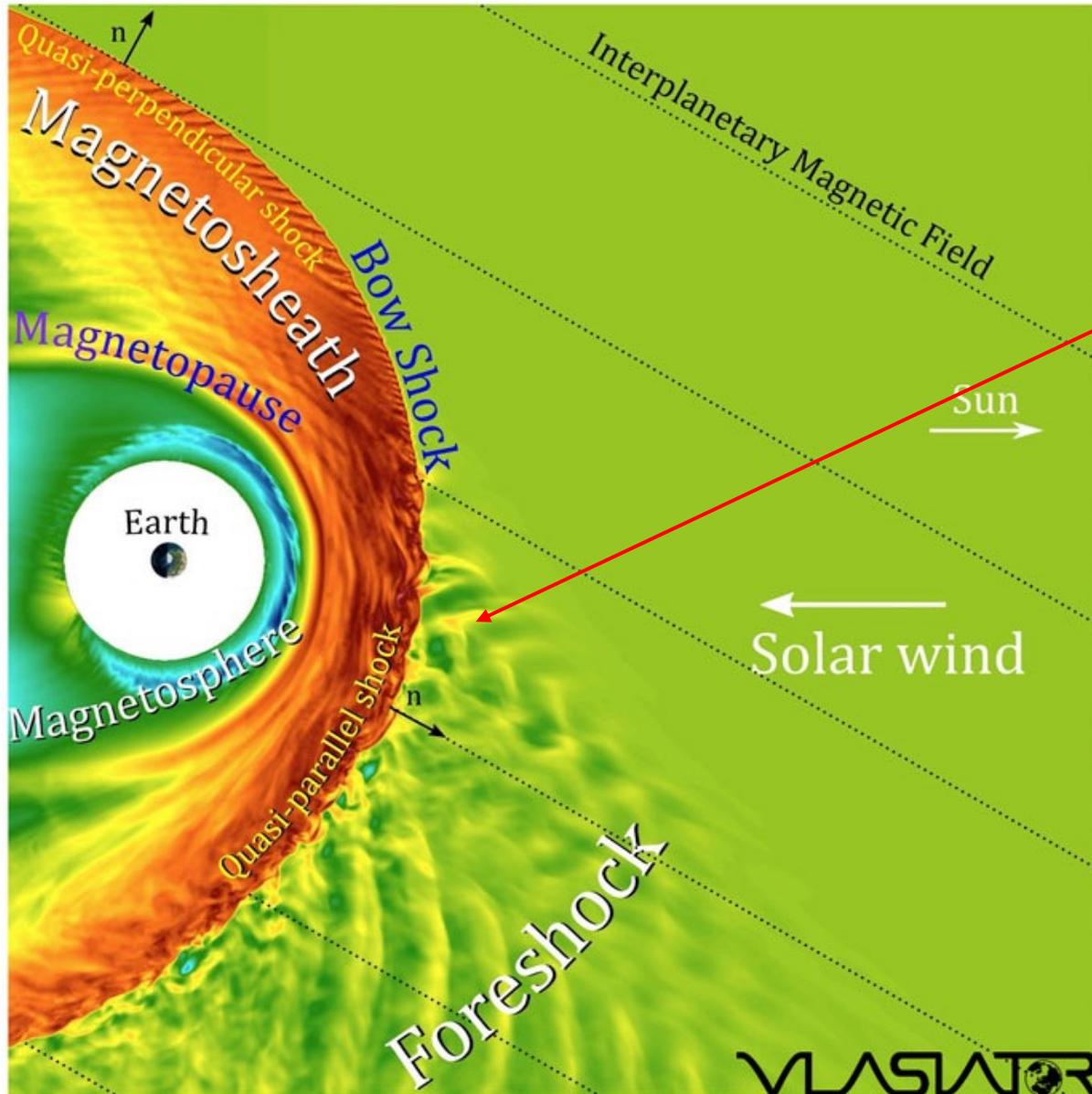
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# Earth's Qpar bow shock and foreshock



Qpar shocks ( $\theta_{Bn} < \sim 45^\circ$ )

- Very **efficient particle accelerators**
- **Transient phenomena** upstream and downstream\*
- ULF waves upstream and downstream
- Kinetic plasma physics
- Wave particle interaction
- Turbulence
- Current sheets & reconnection

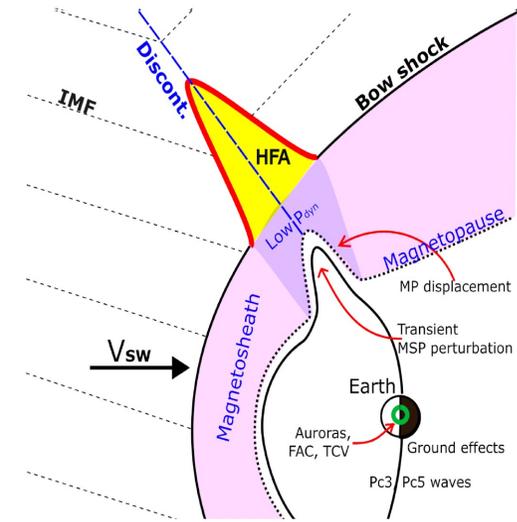
Term to use: Shock-generated transients\*\*



\*\* Credits: Heli Hietala? <https://www.youtube.com/watch?v=SJ88e8rwoDg>

# The anatomy of an HFA

How many: ~several per day!  
How big: ~10s of Re



## Hot Flow Anomalies (HFA)

Kajdič+ (2024)

Discontinuity intersects the bow shock and the convection electric field ( $-V \times B$ ) points towards the sheet on at least one side.

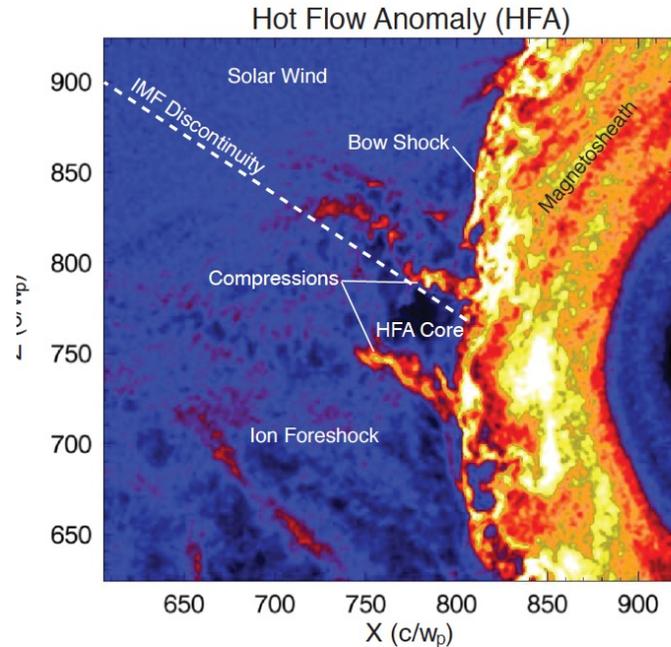


Figure by Nick Omidi

## Crater-like B-field and IMF discontinuity

## Density compressions and cavity

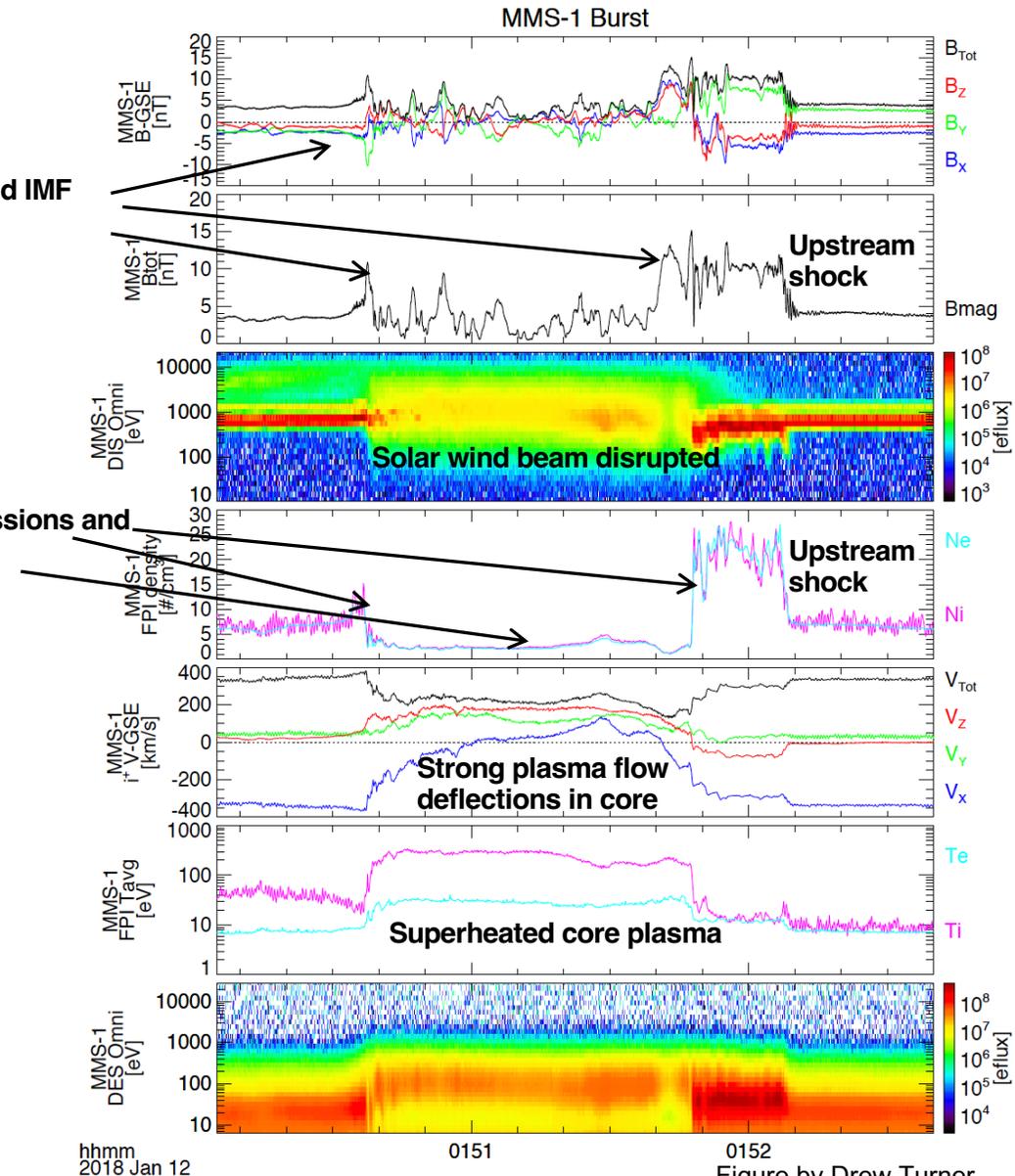
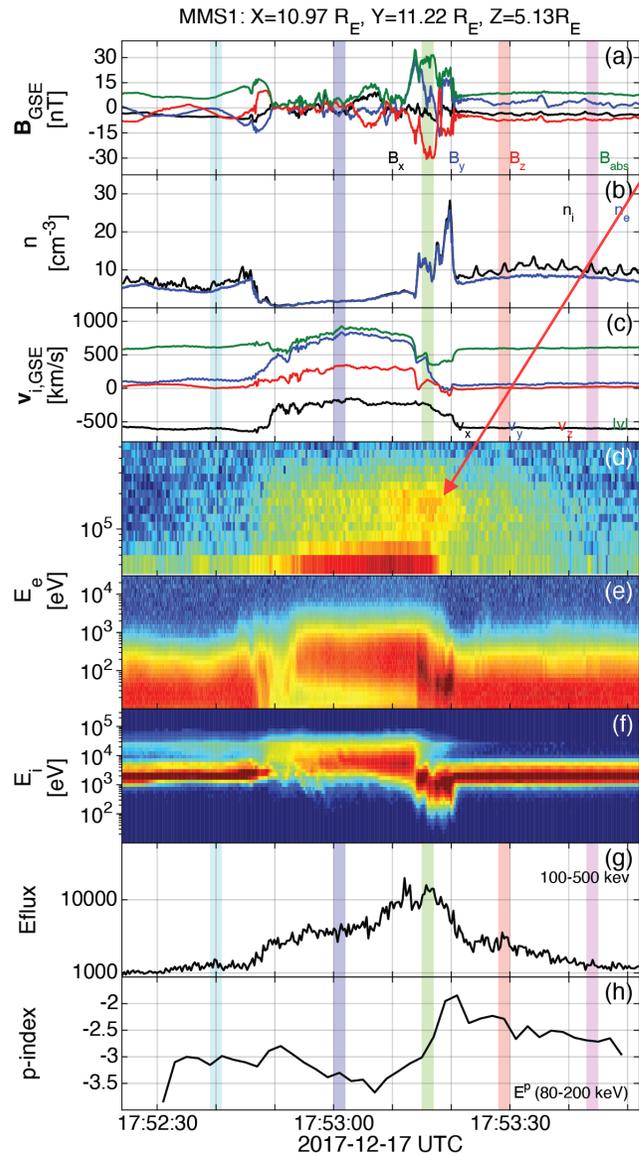


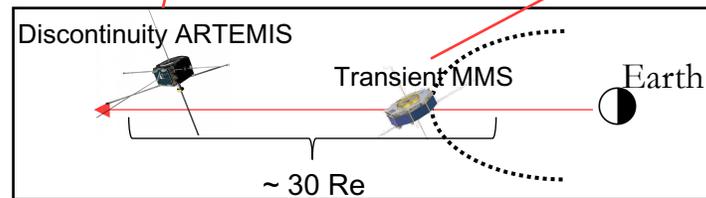
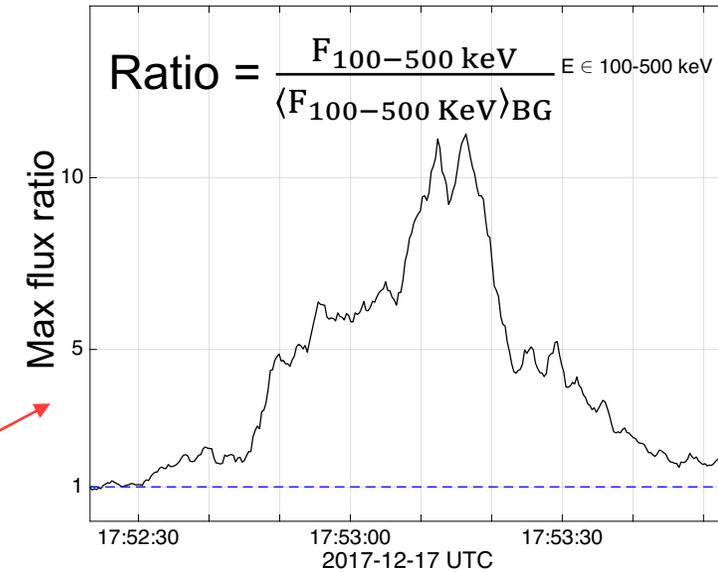
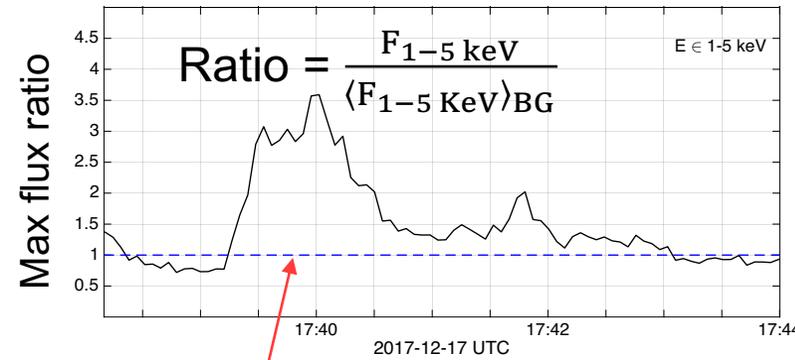
Figure by Drew Turner

# Reinforced Shock Acceleration of Relativistic Electrons



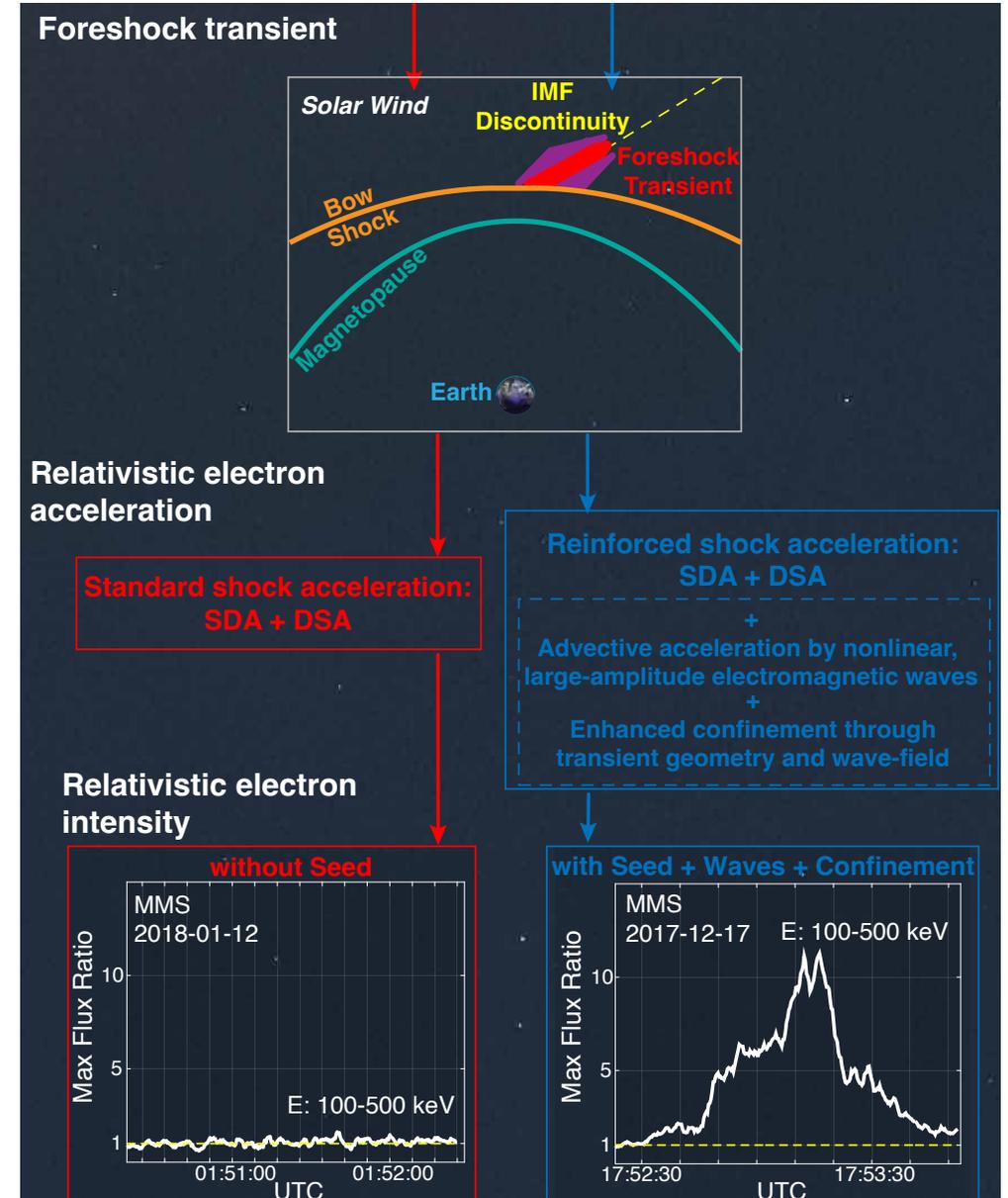
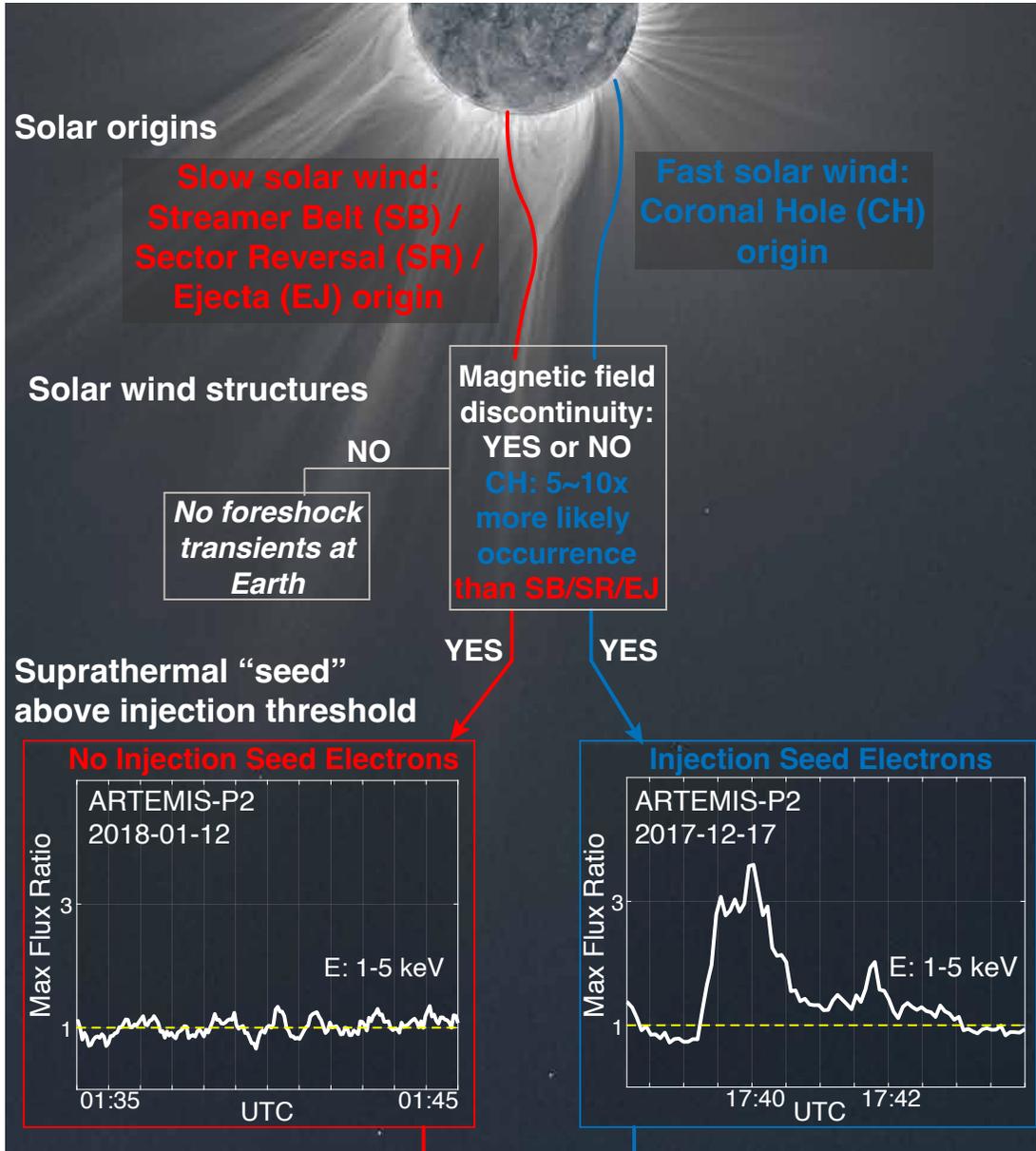
Most energetic electrons observed (**500+keV**) at foreshock transients by MMS showing:

- (a) Seed population from fast solar wind at suprathermal energy range (1-5 KeV)
- (b) Shock acceleration (SDA/DSA) + Wave particle interaction (high frequency “chorus” waves)
- (c) Efficiency factors (Trapping and scattering – geometry with bow shock)



**Key-point:**  
**Seed + Foreshock's Shock + Waves + Efficiency factors = ~MeV electrons before reaching the bow shock.**

# Reinforced Shock Acceleration of Relativistic Electrons



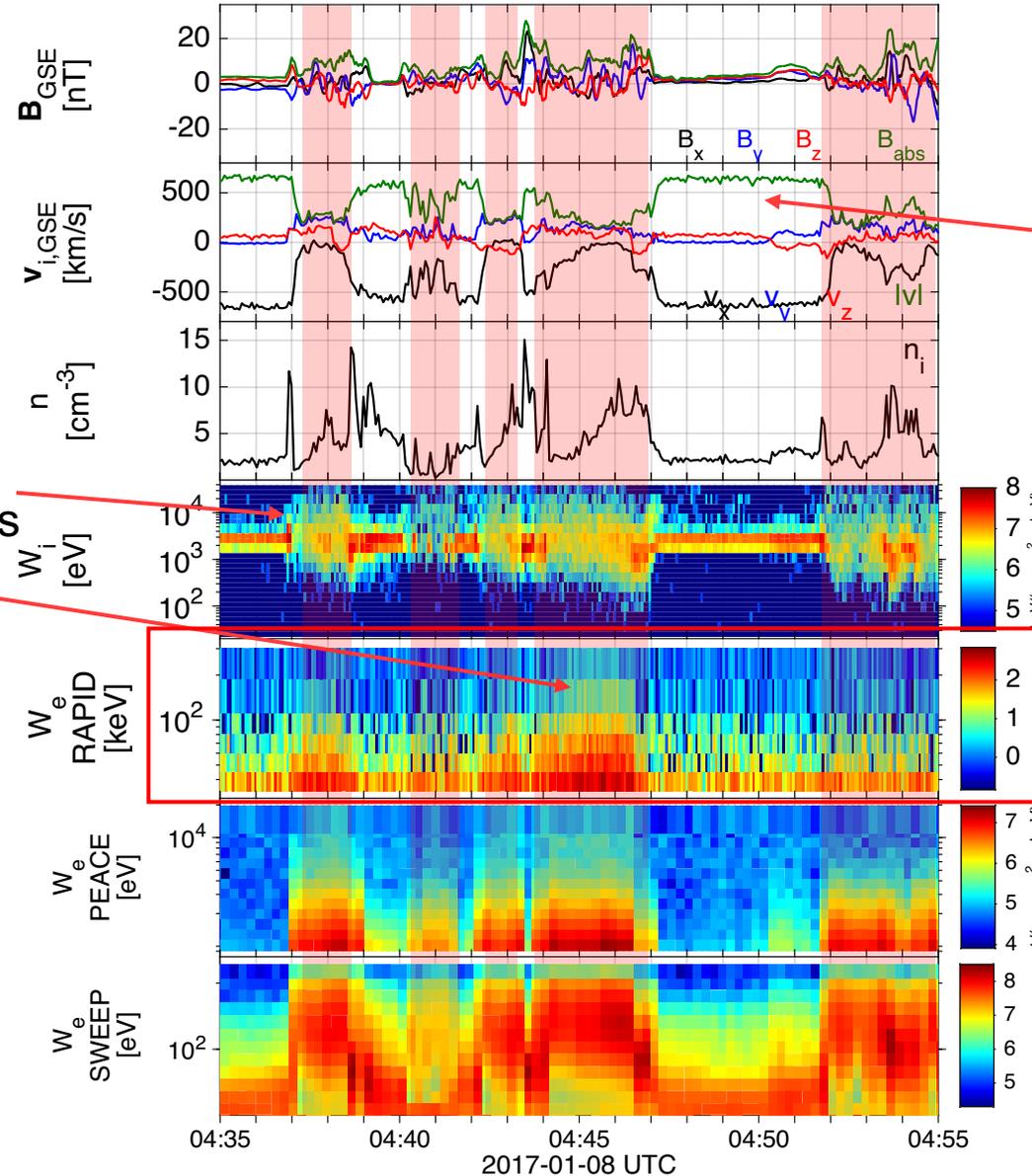
# What happens when we move downstream?

Raptis+ 2024/2025 (to be submitted next week)

*Multi-Mission Observations of Relativistic Electrons and High-Speed Jets  
Linked to Shock Generated Transients*

# Cluster upstream showing foreshock transients

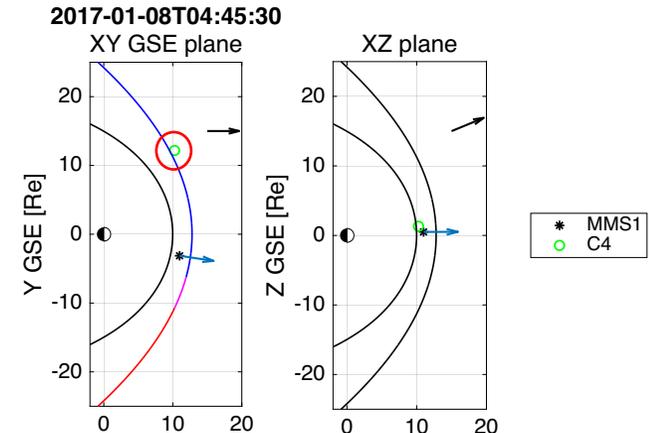
CL 4:  $X=10.27 R_E$ ,  $Y=12.25 R_E$ ,  $Z=-0.45 R_E$



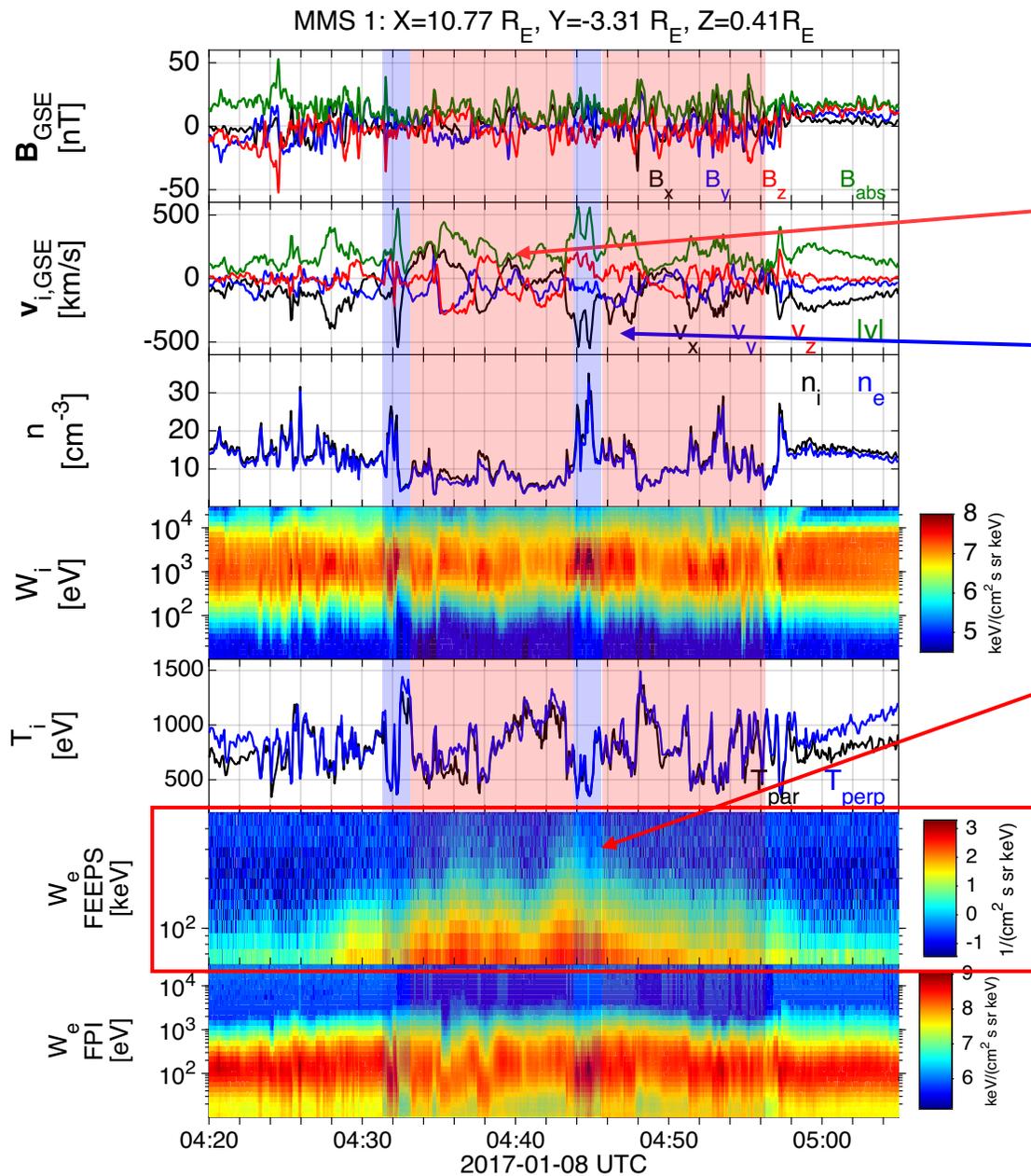
700 km/s Solar Wind!

Fast coronal hole plasma consistent with our previous work ✓

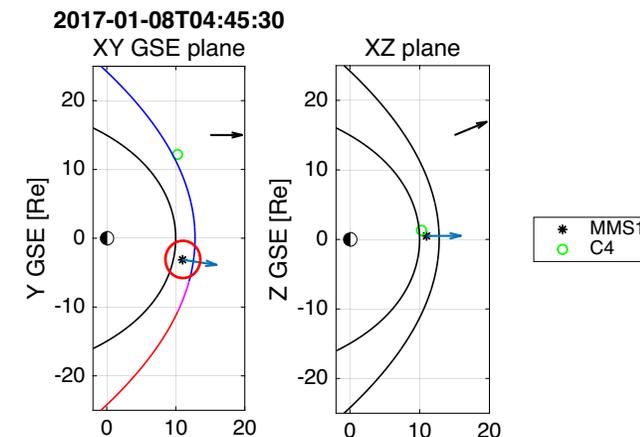
- Several Foreshock Transients
- ~200 keV electrons



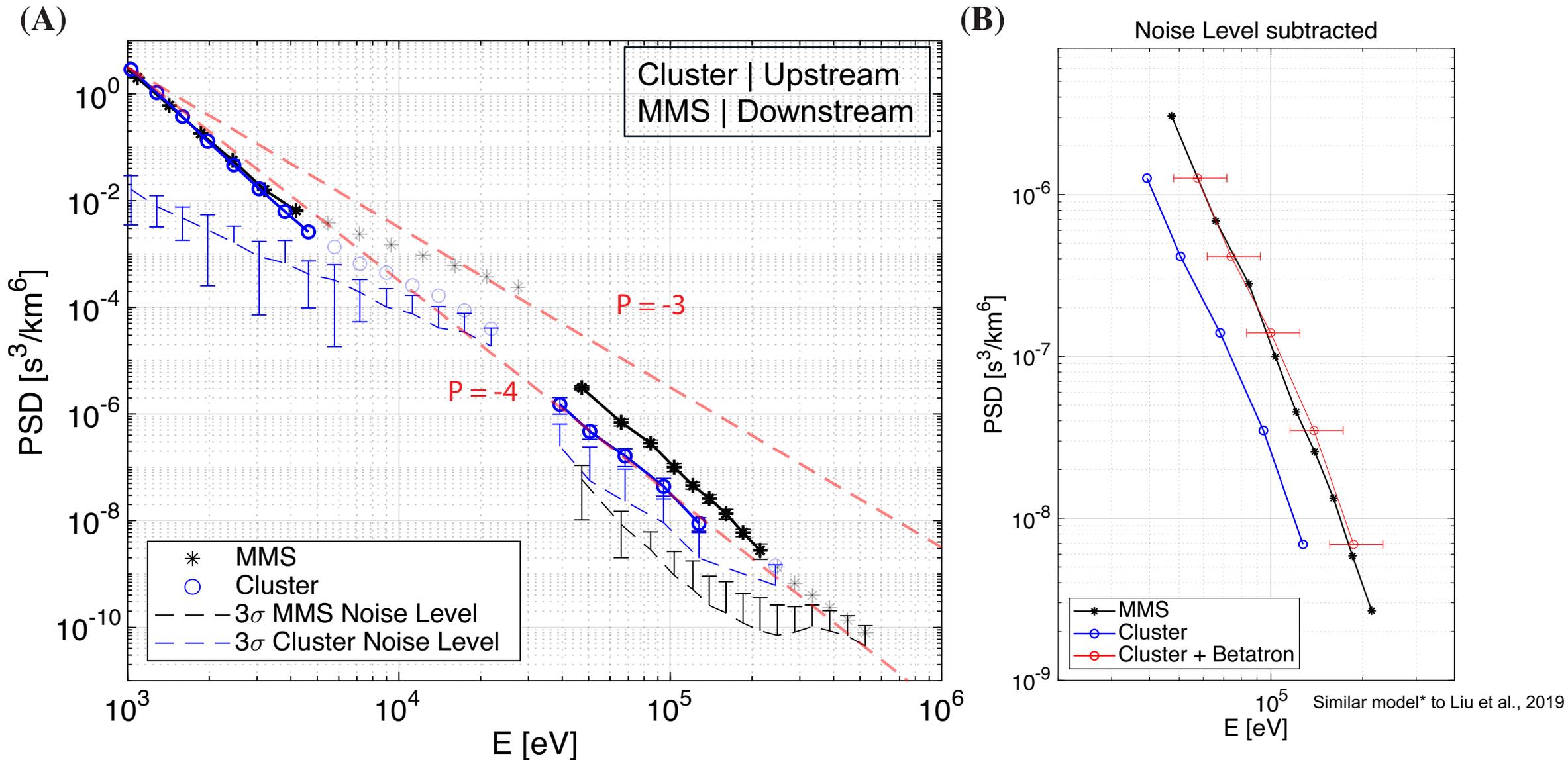
# MMS downstream shows the transmission



- Series of flow anomalies similar to upstream
- Localized jets at edges (both density and velocity peaks)
- Extended  $\sim 300+$  KeV electron interval confined ( $\sim 30$  mins)

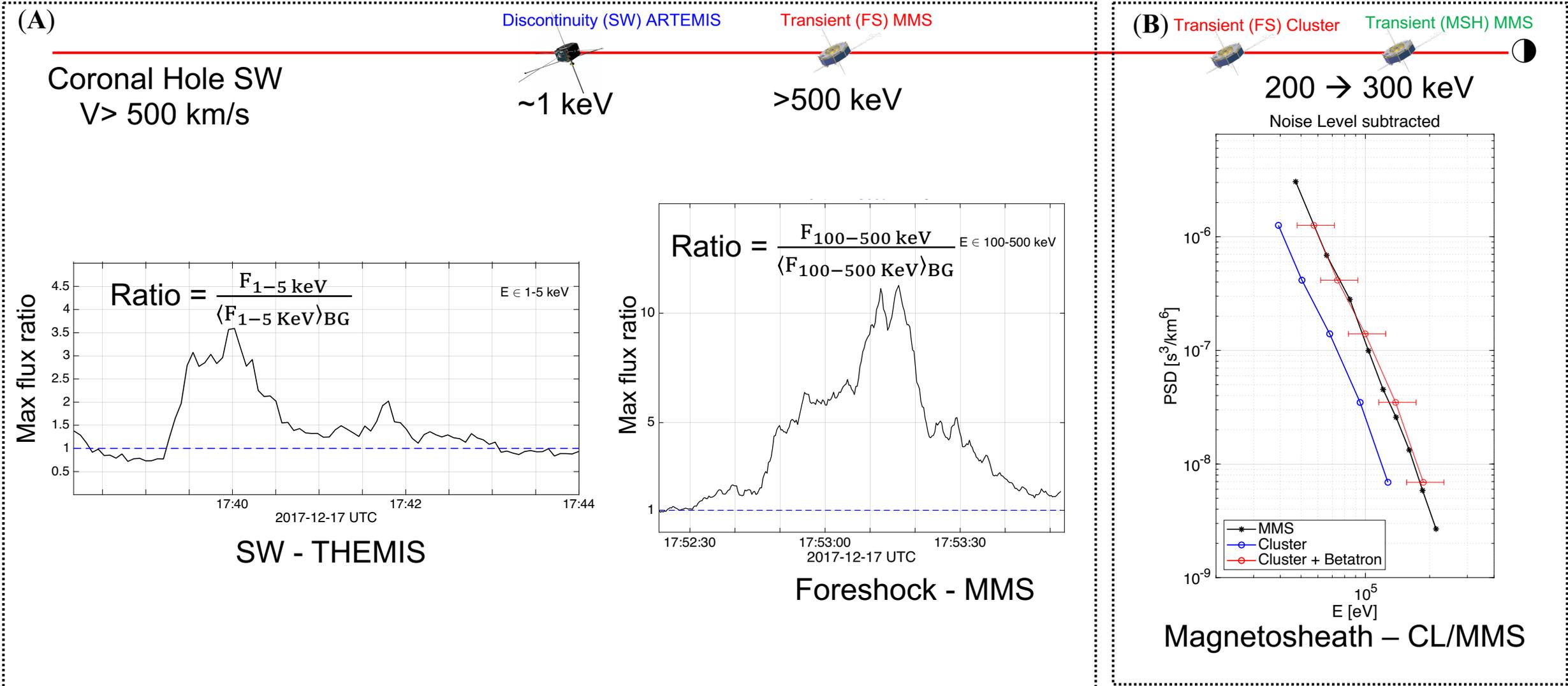


# Electron energization upstream to downstream



\*Betatron in strong scattering  $\rightarrow \frac{\Delta E}{E_0} = \frac{2}{3} \frac{\Delta B}{B_0}$ , where  $\Delta X = X_{\text{down}} - X_{\text{up}}$

# Summary: Electron Multiscale Acceleration



(A) : Raptis+ 2024 (“Accepted”) - *Revealing an Unexpectedly Low Electron Injection Threshold via Reinforced Shock Acceleration*

(B) : Raptis+ 2025 (to be submitted) - *Multi-Mission Observations of Relativistic Electrons and High-Speed Jets Linked to Shock Generated Transients*

# Key take away

(A) : Raptis+ 2024 (Accepted) - *Revealing an Unexpectedly Low Electron Injection Threshold via Reinforced Shock Acceleration*

(B) : Raptis+ 2025 (Ongoing) - *Multi-Mission Observations of Relativistic Electrons and High-Speed Jets Linked to Shock Generated Transients*

## (A) Raptis+ 2024 (“Accepted”)

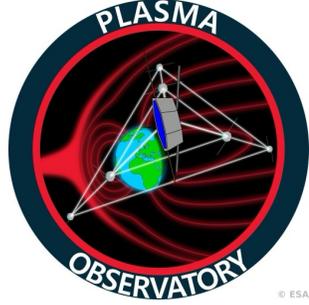
- 1. Revealed that electron injection threshold to obtain relativistic electrons ( $> 500$  keV) at Earth's foreshock can be at the low suprathermal range of  $\sim 1-5$  KeV**
- 2. This seed population is a systematic feature of the fast coronal hole solar wind plasma.**

## (B) Raptis+ 2025 (Ongoing)

- 1. Energized electrons ( $\sim 200$  keV) get further accelerated as transients get transmitted ( $\sim 300$  keV)**
- 2. Transmitted Shock-Generated transients produce localized high-speed jets**

*Strongest electron acceleration may not happen at the shock transition  
Transient phenomena are making shocks much more efficient accelerators*

# Final words



As written on our latest ongoing draft:

“The **multi-mission observations** presented here **emphasize the importance of adopting a multi-scale approach to address shock-related phenomena** and reinforce the necessity of obtaining such data.... proposed missions such as the **Plasma Observatory (PO) mission will be critical to advance our understanding of collisionless shock physics and associated particle energization.** “

What ChatGPT thinks PO will do →

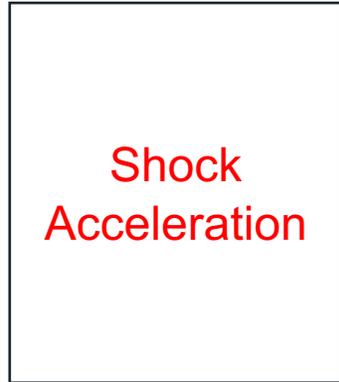
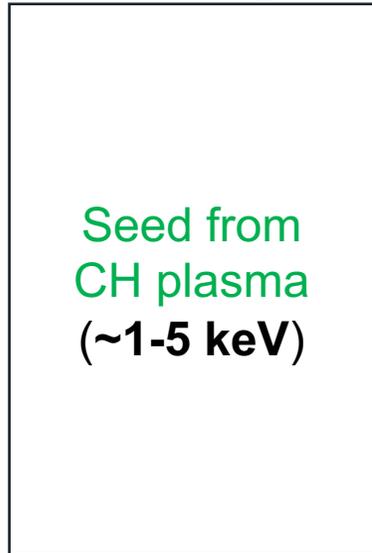


# Extras

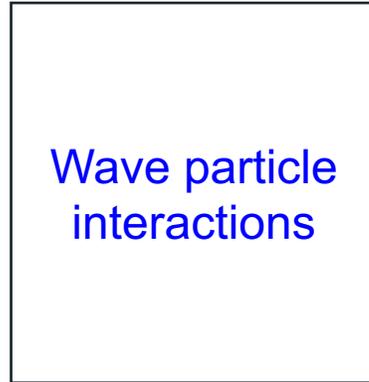
# Reinforced Shock Acceleration of Relativistic Electrons

## Electron acceleration at Foreshock Transient

Injection Threshold  
"Seeding"



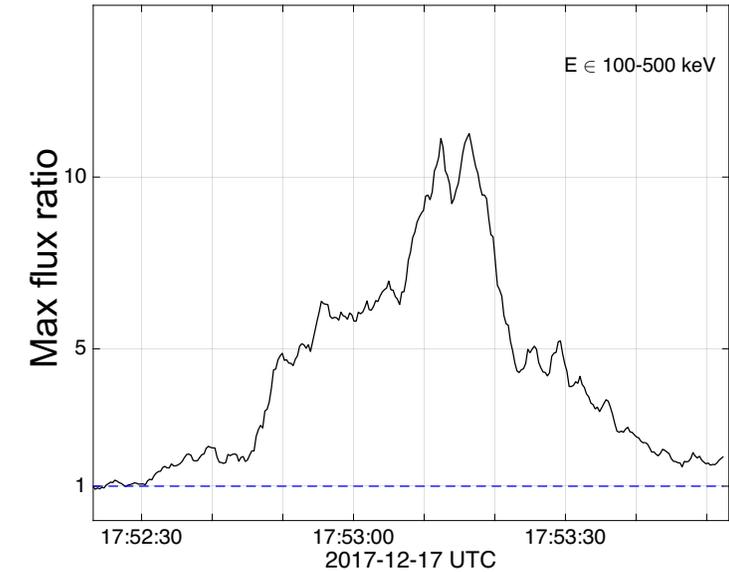
+



Efficient PA scattering & confinement:

- LF whistler waves (Shock Transition)
- Wavefield of HFA's core ( $\Delta B \gg B$ )
- "Magnetic bottle" between edges of HFA and Earth's bow shock (geometry)
- Betatron (electron acceleration, temperature anisotropy)

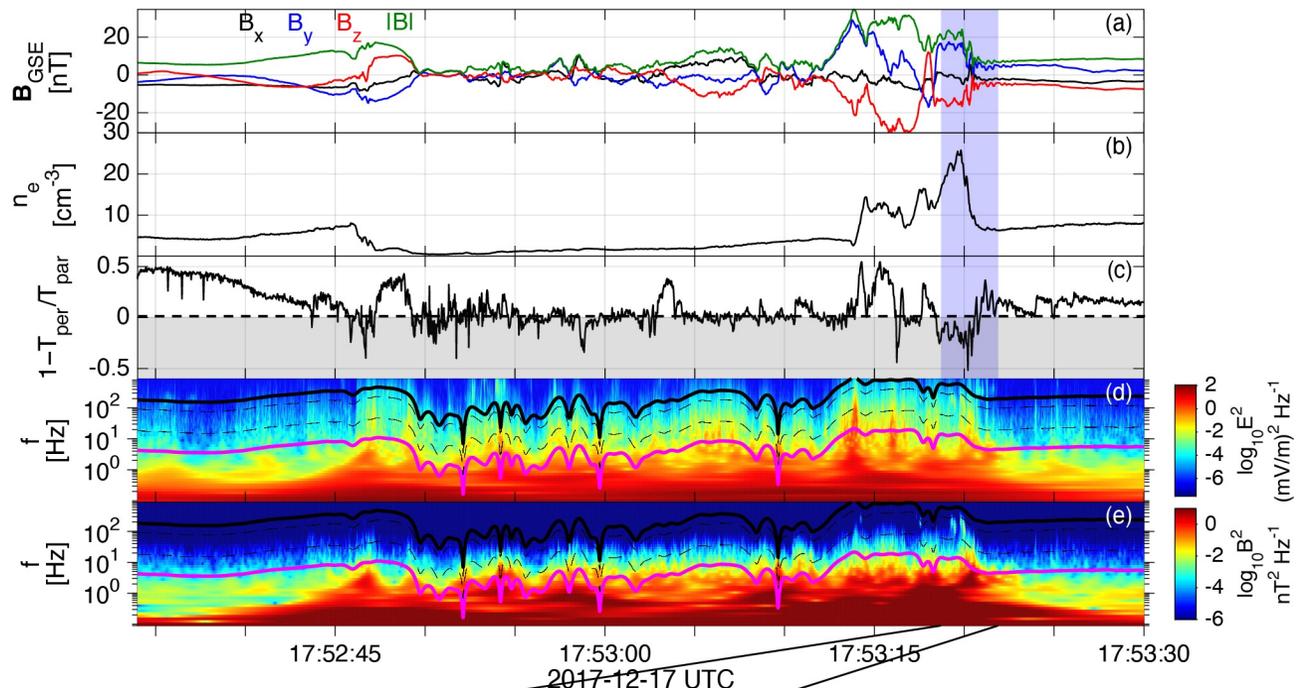
$$\text{Ratio} = \frac{F_{100-500 \text{ KeV}}}{\langle F_{100-500 \text{ KeV}} \rangle_{\text{BG}}}$$



**Key-point:**

**Seed + Foreshock's Shock + Waves + Efficiency factors = ~MeV electrons before reaching the bow shock.**

# Wave analysis of event



## Acceleration:

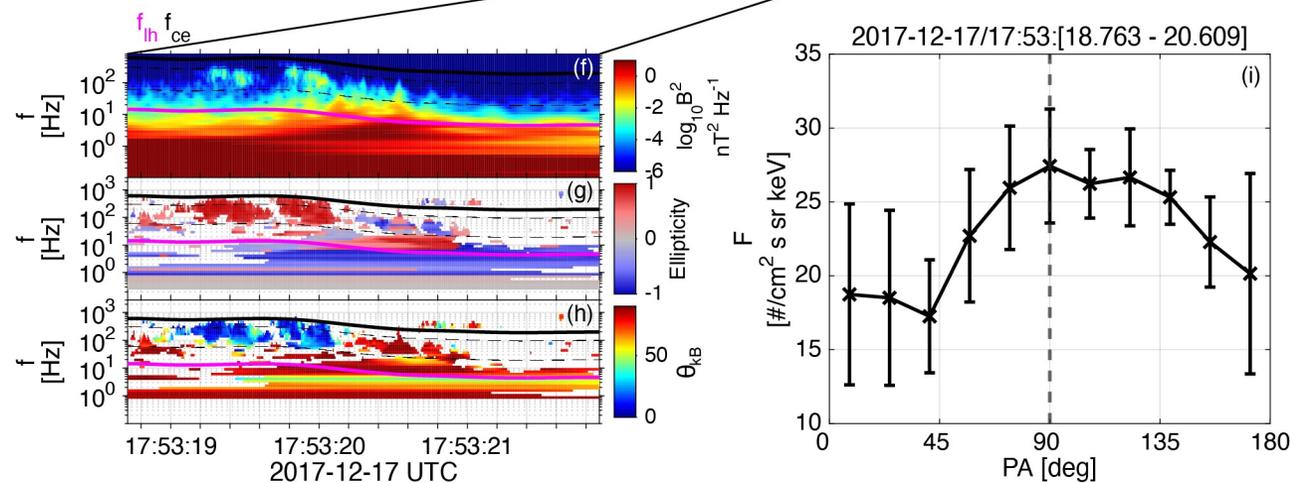
- Shock Acceleration
- Betatron  $\rightarrow$  Temperature anisotropy
- High amplitude electron whistler waves (Chorus) resonance

## Scattering & Trapping:

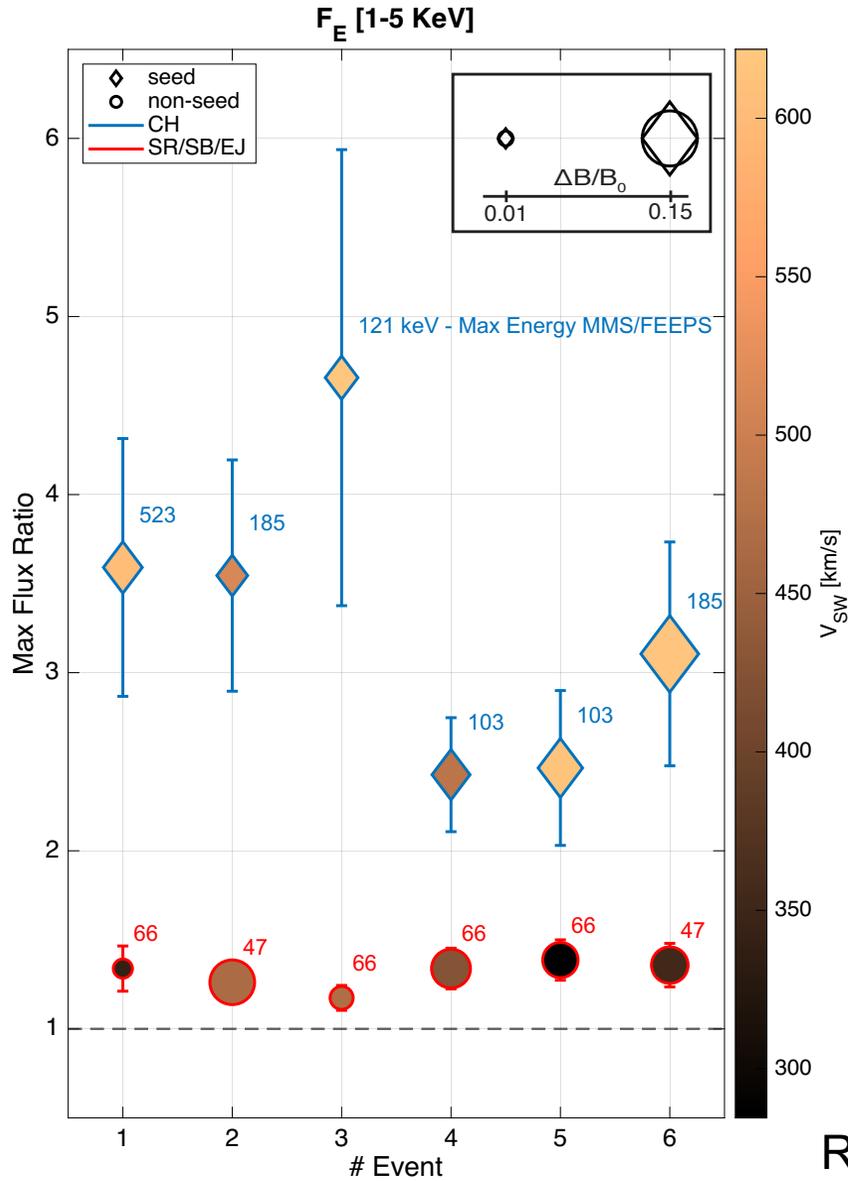
- LF whistler waves (Shock)
- Wavefield of HFA's core
- Geometry of HFA with respect to Earth's bow shock

## Particularly efficient acceleration:

$$\frac{U_E}{U_I} = \frac{\text{Energy density suprathermal electrons}}{\text{Energy density Solar Wind}} \sim 5\%$$



# Statistics for seed and acceleration



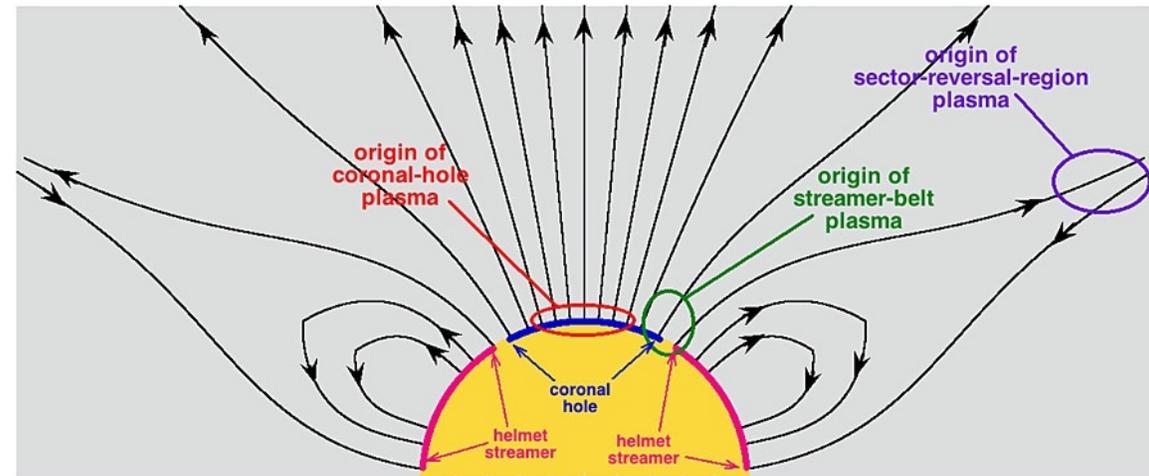
Seed events (FEEPS > 100 KeV)

CH (78%), CH (40%), CH(79%), CH (39%), CH(77%), CH(79%)

No-seed (FEEPS = Noise)

SB (80%), SB(42%), EJ (55%) SB (44%), SR(63%), SB (46%)

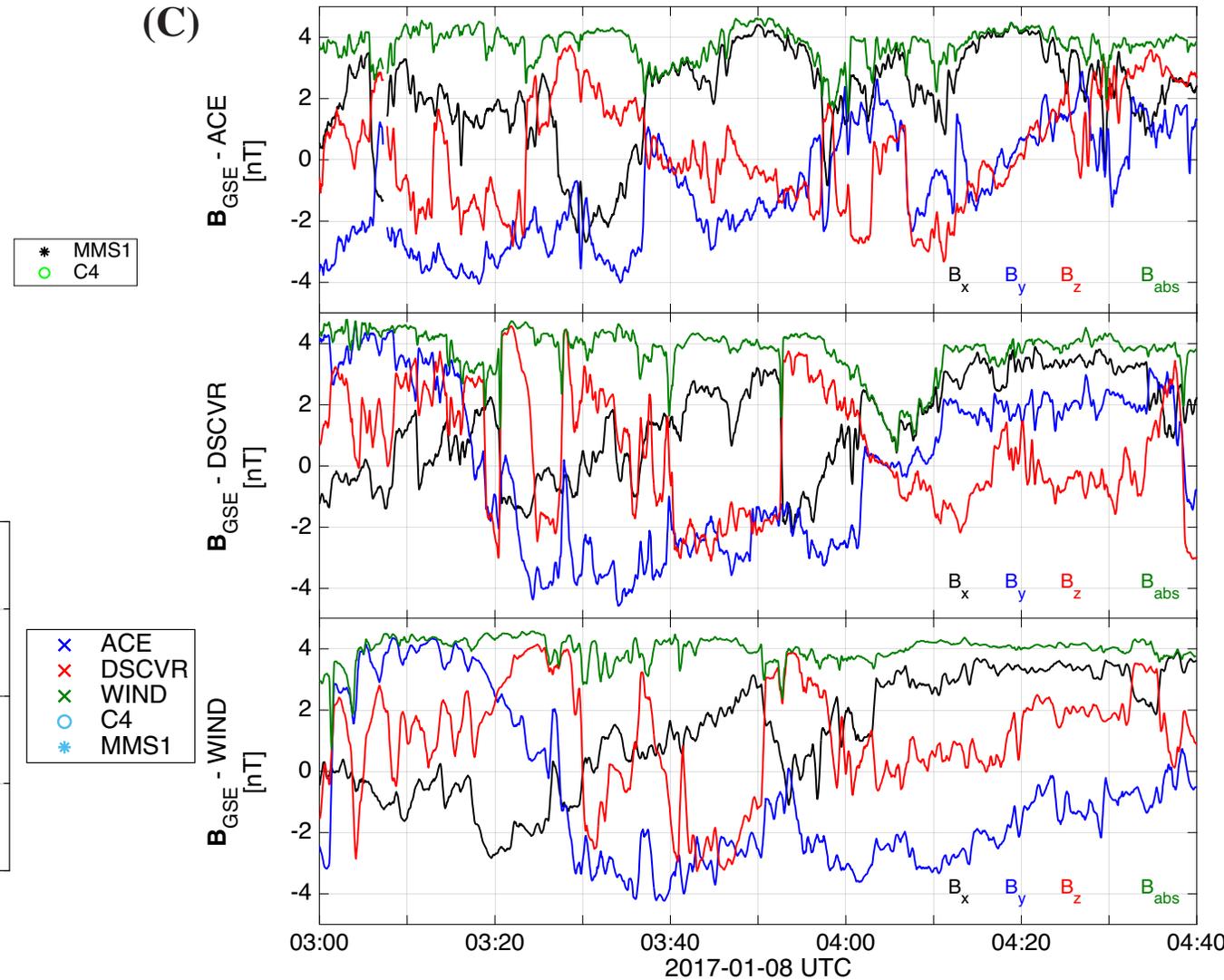
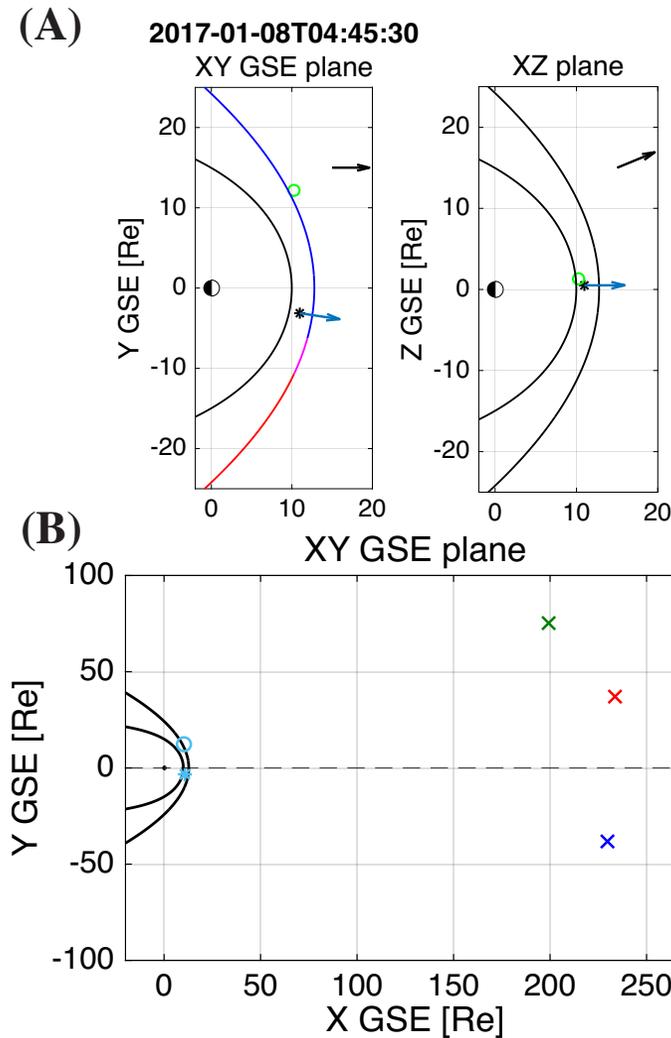
Using: [Xu and Borosvky 2014, Camporeale+ 2017] methodology



$$\text{Ratio} = \frac{F_{1-5 \text{ KeV}}}{\langle F_{1-5 \text{ KeV}} \rangle_{BG}}$$

# Location of SCs and IMF conditions

- (A)**
- Cluster: Upstream
  - MMS : Downstream
- (B)**
- L1 monitors apart
- (C)**
- Temporal & spatial Variability ++
  - Discontinuities ++



# Discussing new models

HOW STANDARDS PROLIFERATE:  
(SEE: A/C CHARGERS, CHARACTER ENCODINGS, INSTANT MESSAGING, ETC\*)

